

RECEIVED: April 3, 2015

REVISED: August 9, 2015

ACCEPTED: September 8, 2015

PUBLISHED: October 5, 2015

Group manifold approach to higher spin theory

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ABSTRACT: We consider the group manifold approach to higher spin theory. The deformed local higher spin transformation is realized as the diffeomorphism transformation in the group manifold \mathbf{M} . With the suitable rheonomy condition and the torsion constraint imposed, the unfolded equation can be obtained from the Bianchi identity, by solving which, fields in \mathbf{M} are determined by the multiplet at one point, or equivalently, by $(W_\mu^{[a(s-1),b(0)]}, H)$ in $\text{AdS}_4 \subset \mathbf{M}$. Although the space is extended to \mathbf{M} to get the geometrical formulation, the dynamical degrees of freedom are still in AdS_4 . The $4d$ equations of motion for $(W_\mu^{[a(s-1),b(0)]}, H)$ are obtained by plugging the rheonomy condition into the Bianchi identity. The proper rheonomy condition allowing for the maximum on-shell degrees of freedom is given by Vasiliev equation. We also discuss the theory with the global higher spin symmetry, which is in parallel with the WZ model in supersymmetry.

KEYWORDS: Higher Spin Gravity, Higher Spin Symmetry

ARXIV EPRINT: [1501.02322](https://arxiv.org/abs/1501.02322)

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1 Introduction

Group manifold approach provides a natural geometrical formulation for supergravity [1–4]. The starting point is the supergroup $\overline{\text{Osp}}(1/4)$ or $\text{Osp}(1/4)$. Supergravity field and matter field are vielbein 1-form $\nu_{\bar{M}}^A$ and 0-form H on the group manifold \mathbf{M} , $A, \bar{M} = 1, \dots, \dim \overline{\text{Osp}}(1/4)$. Local super Poincaré transformation is realized as the diffeomorphism transformation on \mathbf{M} . The curvature $R_{\bar{M}\bar{N}}^A$ for the 1-form can be defined, on which, the rheonomy condition is imposed [1–4]. The condition requires that $R_{\bar{M}\bar{N}}^A$ can be algebraically expressed in terms of its purely “inner” components $R_{\mu\nu}^A$ with $\mu, \nu = 1, 2, 3, 4$ the indices in a four-dimensional submanifold M_4 . Namely,

$$R_{\bar{M}\bar{N}}^A = r_{\bar{M}\bar{N}}^A|_B^{\mu\nu} R_{\mu\nu}^B, \quad \text{or} \quad R_{CD}^A = r_{CD}^A|_B^{ab} R_{ab}^B \quad (1.1)$$

where $r_{\bar{M}\bar{N}}^A|_B^{\mu\nu}$ and $r_{CD}^A|_B^{ab}$ are constant holonomic and anholonomic tensors. $a, b = 1, 2, 3, 4$. The rheonomy condition ensures that fields on the whole \mathbf{M} are determined by fields on M_4 . So the final dynamics is still in M_4 , where the diffeomorphism transformation reduces

to the on-shell super Poincaré transformation of the $4d$ fields. The equations of motion in M_4 are obtained by plugging the rheonomy condition into the Bianchi identity. Instead of imposing the rheonomy condition, one can also construct the extended action, which is the integration of some 4-form on a $4d$ submanifold M_4 . Variation of the action with respect to both fields and M_4 gives the rheonomy condition as well as the $4d$ equations of motion.

In this paper, we will reformulate the group manifold method, adding an infinite number of auxiliary fields so that the final system is equivalent to the unfolded dynamics approach which is convenient for higher spin theory [5]. For simplicity, we will consider the minimal bosonic $4d$ HS algebra $\text{ho}(1|2 : [3, 2])$ with spin $s = 0, 2, \dots$ [6]. The corresponding group manifold is denoted as \mathbf{M} . Fields are 1-form W_M^α and 0-form H on \mathbf{M} with the curvature 2-form

$$dW^\alpha = \frac{1}{2}(f_{\beta\gamma}^\alpha + R_{\beta\gamma}^\alpha)W^\beta \wedge W^\gamma = \frac{1}{2}\hat{f}_{\beta\gamma}^\alpha W^\beta \wedge W^\gamma \quad (1.2)$$

and the 1-form

$$dH = H_\alpha W^\alpha. \quad (1.3)$$

$\bar{M} = 1, 2, \dots, \dim \text{ho}(1|2 : [3, 2])$. $\alpha \sim [a(s-1), b(t)]$ is in the adjoint representation of $\text{ho}(1|2 : [3, 2])$. $f_{\beta\gamma}^\alpha$ is the structure constant of $\text{ho}(1|2 : [3, 2])$. The deformed higher spin transformation is the diffeomorphism transformation on \mathbf{M} .

The rheonomy condition is

$$\begin{aligned} \hat{f}_{\beta\gamma}^\alpha &= \hat{f}_{\beta\gamma}^\alpha(R_{ab}^{[a(s-1), b(s-1)]}, R_{ab; c_1}^{[a(s-1), b(s-1)]}, \dots, H, H_{c_1}, \dots) \\ H_\alpha &= h_\alpha(R_{ab}^{[a(s-1), b(s-1)]}, R_{ab; c_1}^{[a(s-1), b(s-1)]}, \dots, H, H_{c_1}, \dots), \end{aligned} \quad (1.4)$$

where

$$R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]} = \partial_{c_n} \dots \partial_{c_1} R_{ab}^{[a(s-1), b(s-1)]}, \quad H_{c_1 \dots c_n} = \partial_{c_n} \dots \partial_{c_1} H. \quad (1.5)$$

$\partial_c = W_c^{\bar{M}} \partial_{\bar{M}}$, $a, b, c = 1, 2, 3, 4$. a is the abbreviation for the $[0, a]$ element of $\text{ho}(1|2 : [3, 2])$. Different from the supergravity situation, the curvature depends on the “inner” components as well as their “inner” derivatives. This is the most generic rheonomy condition. (1.4) together with the Bianchi identity gives the unfolded equation

$$\begin{aligned} dW^\alpha &= \frac{1}{2}\hat{f}_{\beta\gamma}^\alpha W^\beta \wedge W^\gamma, \\ dR_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]} &= r_{ab; c_1 \dots c_n \gamma}^{[a(s-1), b(s-1)]} W^\gamma, \\ dH_{c_1 \dots c_n} &= h_{c_1 \dots c_n \gamma} W^\gamma, \end{aligned} \quad (1.6)$$

from which, $(W^\alpha, R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ on the whole \mathbf{M} is determined by its value at one point. On $\text{AdS}_4 \subset \mathbf{M}$, we have the further relation

$$\begin{aligned} &(R_{ab}^{[a(s-1), b(s-1)]}, R_{ab; c_1}^{[a(s-1), b(s-1)]}, \dots, H, H_{c_1}, \dots) \\ &\sim (W_\mu^{[a(s-1), b(0)]}, \partial_{\nu_1} W_\mu^{[a(s-1), b(0)]}, \dots, H, \partial_{\nu_1} H, \dots), \end{aligned} \quad (1.7)$$

where ∂_μ is the derivative on AdS_4 . So equivalently, with $(W_\mu^{[a(s-1),b(0)]}, H)$ given on AdS_4 , (W_M^α, H) on the whole \mathbf{M} can be determined up to a gauge transformation. The dynamical 1-form fields are $W^{[a(s-1),b(0)]}$, which is because in (1.4), the torsion constraint is also implicitly imposed: $\hat{f}_{\beta\gamma}^\alpha$ and H_α do not depend on $R_{ab;c_1\cdots c_n}^{[a(s-1),b(t)]}$ with $t \neq s-1$. For 0-form, the deformed higher spin transformation is $\xi^{\bar{M}}\partial_{\bar{M}} = \epsilon^\alpha\partial_\alpha$, under which, the multiplet $(R_{ab}^{[a(s-1),b(s-1)]}, R_{ab;c_1}^{[a(s-1),b(s-1)]}, \dots, H, H_{c_1}, \dots)$ forms the complete representation on-shell.

The whole dynamics is encoded in functions $(\hat{f}_{\beta\gamma}^\alpha, h_\alpha)$, which should satisfy the Bianchi identity and also give the correct free theory limit. With the unfolded equation plugged in, the Bianchi identities are polynomials of $(R_{ab;c_1\cdots c_n}^{[a(s-1),b(s-1)]}, H_{c_1\cdots c_n})$, by solving which, $(\hat{f}_{\beta\gamma}^\alpha, h_\alpha)$ is determined with the rest constraints on $(R_{ab;c_1\cdots c_n}^{[a(s-1),b(s-1)]}, H_{c_1\cdots c_n})$ acting as the $4d$ equations of motion. The procedure is simple in supergravity but is extremely complicated in higher spin theory. Instead of fixing $(\hat{f}_{\beta\gamma}^\alpha, h_\alpha)$ and getting the $4d$ equations of motion by solving the Bianchi identity, one can first identify the on-shell degrees of freedom, for example, $\Phi^{\tilde{\sigma}} \sim \Phi^{[a(s+n),b(s)]}$ in the twisted-adjoint representation of the higher spin algebra, then find the suitable $(\hat{f}_{\beta\gamma}^\alpha, h_\alpha)$ so that the Bianchi identity is satisfied for the arbitrary $\Phi^{\tilde{\sigma}}$.

$$\{H_{c_1\cdots c_n}, n = 0, 1, \dots\} \cup \{R_{ab;c_1\cdots c_n}^{[a(s-1),b(s-1)]}, s = 2, 4, \dots, n = 0, 1, \dots\} \quad (1.8)$$

and $\{\Phi^{[a(s+n),b(s)]}, s = 0, 2, \dots, n = 0, 1, \dots\}$ have the same number of indices. With the $4d$ equations of motion imposed on (1.8), the two may contain the same number of degrees of freedom. Written in terms of $\Phi^{\tilde{\sigma}}$, the unfolded equation becomes

$$dW^\alpha = \frac{1}{2}\bar{f}_{\beta\gamma}^\alpha(\Phi^{\tilde{\sigma}})W^\beta \wedge W^\gamma, \quad d\Phi^{\tilde{\alpha}} = F_{\beta}^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}})W^\beta. \quad (1.9)$$

It remains to find $(\bar{f}_{\beta\gamma}^\alpha, F_{\beta}^{\tilde{\alpha}})$ satisfying the Bianchi identity and also giving rise to the correct free theory limit.¹ Vasiliev theory gives the elegant solution to this problem [7–9]. By solving the Z part of the Vasiliev equation order by order, one may finally get the required $(\bar{f}_{\beta\gamma}^\alpha, F_{\beta}^{\tilde{\alpha}})$ [10].

For supersymmetry, it is also possible to study the dynamics of the 0-form matter on group manifold with the fixed background such as the WZ model. The component expansion of the 0-form on superspace gives the spin 0 and 1/2 fields in $4d$. The allowed gauge transformation is the global super Poincaré transformation, which is the diffeomorphism transformation on \mathbf{M} preserving the background. For higher spin theory, one can similarly consider the 0-form H on \mathbf{M} with

$$dW_0^\alpha = \frac{1}{2}f_{\beta\gamma}^\alpha W_0^\beta \wedge W_0^\gamma, \quad dH = H_\alpha W_0^\alpha. \quad (1.10)$$

W_0^α describes the background with the vanishing curvature. The system has the global HS symmetry. The component expansion of H on \mathbf{M} gives the spin $s = 0, 2, \dots$ fields $R_{a_1\cdots a_s, b_1\cdots b_s}^s$. On the other hand, the linearized Vasiliev equation for the 0-forms on background W_0^α is

$$d\Phi^{\tilde{\alpha}} = k_{\beta\gamma}^{\tilde{\alpha}}\Phi^{\tilde{\gamma}}W_0^\beta, \quad (1.11)$$

¹As is shown in appendix C, there are $(\bar{f}_{\beta\gamma}^\alpha, F_{\beta}^{\tilde{\alpha}})$ satisfying the Bianchi identity but failing to give the correct free theory limit. It is unclear whether the two requirements can uniquely fix $(\bar{f}_{\beta\gamma}^\alpha, F_{\beta}^{\tilde{\alpha}})$ (up to a field redefinition) or not.

which is also invariant under the global HS transformation. $k_{\beta\tilde{\gamma}}^{\tilde{\alpha}}$ is the constant. With $\Phi \equiv \Phi^{[a(0),b(0)]} = H$, from $d\Phi = k_{\beta\tilde{\gamma}}\Phi^{\tilde{\gamma}}W_0^\beta$, we have $H_\beta = k_{\beta\tilde{\gamma}}\Phi^{\tilde{\gamma}}$. $R_{a_1\dots a_s,b_1\dots b_s}^s$ can then be taken as the Weyl tensor of the linearized HS theory. With the space extended from AdS_4 to \mathbf{M} , 0-forms in the linearized Vasiliev theory get the interpretation as the derivatives of a single 0-form H on \mathbf{M} .

The rest of the paper is organized as follows. In section 2, we construct a symmetric space M with the higher spin transformation group the isometry group. In section 3, we consider the theory with the local higher spin symmetry. The discussion and conclusion are given in section 4.

2 Symmetric space from the higher spin algebra

We will consider the minimal bosonic higher spin theory in AdS_4 with the coordinate w^μ , $\mu = 1, 2, 3, 4$. The related HS algebra is $\text{ho}(1|2 : [3, 2])$ with the basis $\{t_\alpha \sim t_{A_1\dots A_{s-1}, B_1\dots B_{s-1}}\}$ [6]. $t_{A_1\dots A_{s-1}, B_1\dots B_{s-1}}$ is in irreducible representations of $\text{SO}(3, 2)$ characterized by two row rectangular Young tableaux, $A_i, B_i = 0, 1, 2, 3, 4$, $s = 2, 4, \dots$

$$\begin{aligned} t_{A_1\dots A_{s-1}, B_1\dots B_{s-1}} &= t_{\{A_1\dots A_{s-1}\}, B_1\dots B_{s-1}} = t_{A_1\dots A_{s-1}, \{B_1\dots B_{s-1}\}}, \\ t_{\{A_1\dots A_{s-1}, A_s\} B_2\dots B_{s-1}} &= 0, \quad t_{A_1\dots A_{s-3} CC, B_1\dots B_{s-1}} = 0. \end{aligned} \quad (2.1)$$

With $a_i, b_i = 1, 2, 3, 4$, basis of $\text{ho}(1|2 : [3, 2])$ can be rewritten as

$$\begin{aligned} \{t_\alpha\} &= \{t_{A_1\dots A_{s-1}, B_1\dots B_{s-1}}\} \\ &= \{t_{0\dots 0, b_1\dots b_{s-1}}, t_{0\dots 0a_1, b_1\dots b_{s-1}}, t_{0\dots 0a_1a_2, b_1\dots b_{s-1}}, \dots, t_{a_1\dots a_{s-1}, b_1\dots b_{s-1}}\}. \end{aligned} \quad (2.2)$$

Let

$$\{t_Q\} = \{t_{0\dots 0a_1, b_1\dots b_{s-1}}, t_{0\dots 0a_1a_2a_3, b_1\dots b_{s-1}}, \dots, t_{a_1\dots a_{s-1}, b_1\dots b_{s-1}}\} \quad (2.3)$$

be the basis of $a[E]$,

$$\{t_A\} = \{t_{0\dots 0, b_1\dots b_{s-1}}, t_{0\dots 0a_1a_2, b_1\dots b_{s-1}}, \dots, t_{0a_1\dots a_{s-2}, b_1\dots b_{s-1}}\} \quad (2.4)$$

be the basis of K , $\text{ho}(1|2 : [3, 2]) = a[E] \oplus K$.

$$[a[E], a[E]] \subset a[E], \quad [a[E], K] \subset K, \quad [K, K] \subset a[E]. \quad (2.5)$$

$a[E]$ is a subalgebra of $\text{ho}(1|2 : [3, 2])$ generating a subgroup E . The coset space $G[\text{ho}(1|2 : [3, 2])]/E$ is a symmetric space according to (2.5). With the group given, it is a standard procedure in mathematics to construct the group manifold \mathbf{M} for $G[\text{ho}(1|2 : [3, 2])]$ ² and the symmetric space M for $G[\text{ho}(1|2 : [3, 2])]/E$. In the following, we will give a construction based on the operators and the conserved charges of the quantum higher spin theory in AdS_4 . For earlier work on space with the tensor coordinates, see [11, 12].

²HS transformation group is the global symmetry group of the $3d$ $O(N)$ vector model and the dual minimal bosonic HS theory in AdS_4 . The related algebra is $\text{ho}(1|2 : [3, 2])$. The group can be non-connected, just as $\text{SO}(3, 1)$. Here $G[\text{ho}(1|2 : [3, 2])]$ refers to the the connected piece containing the identity, which is a simple group. So the related group manifold \mathbf{M} is also connected.

In quantum higher spin theory, there are conserved charges $\{Q_{A_1 \dots A_{s-1}, B_1 \dots B_{s-1}}\}$ in one-to-one correspondence with $\{t_{A_1 \dots A_{s-1}, B_1 \dots B_{s-1}}\}$. In particular, $\{Q_{A_1, B_1}\}$ are generators of $\text{SO}(3, 2)$. Suppose 0 is a point in the bulk of AdS_4 , for example, $(1, 0, 0, 0, 0)$ in $x_0^2 - x_1^2 - x_2^2 - x_3^2 + x_4^2 = 1$, and $O(0)$ is the operator for the spin 0 field at 0, then the orbit generated by $\text{SO}(3, 2)$ gives operators for the spin 0 field in the entire AdS_4 .

$$\{O(u)|u \in \text{AdS}_4\} = \{gO(0)g^{-1}|g \in \text{SO}(3, 2)\}, \quad (2.6)$$

where $g = e^{i\omega^{A_1, B_1} Q_{A_1, B_1}}$. Aside from AdS_4 , the orbit generated by $G[\text{ho}(1|2 : [3, 2])]$ gives operators in an enlarged space M .

$$\{O(z)|z \in M\} = \{gO(0)g^{-1}|g \in G[\text{ho}(1|2 : [3, 2])]\}, \quad (2.7)$$

where $g = e^{i\omega^{A_1 \dots A_{s-1}, B_1 \dots B_{s-1}} Q_{A_1 \dots A_{s-1}, B_1 \dots B_{s-1}}}$. In $G[\text{ho}(1|2 : [3, 2])]$, there is a subgroup $E(z)$, $\forall e \in E(z)$, $eO(z)e^{-1} = O(z)$. The higher spin algebra is decomposed as

$$\text{ho}(1|2 : [3, 2]) = K(z) \oplus a[E(z)] = g(z)K(0)g(z)^{-1} \oplus g(z)a[E(0)]g(z)^{-1} \quad (2.8)$$

with $K(z)$ the tangent space of M at z . M is the coset space $G[\text{ho}(1|2 : [3, 2])]/E$. In particular, $\text{SO}(3, 1) \subset E$, $\text{SO}(3, 2) \subset G[\text{ho}(1|2 : [3, 2])]$, $\text{AdS}_4 = \text{SO}(3, 2)/\text{SO}(3, 1)$, so M has a fiber bundle structure with the fiber AdS_4 attached at each point of the base manifold.

It remains to determine the subalgebra $a[E]$. Although the direct quantization of the higher spin theory in AdS_4 is still not available, its CFT dual is quite simple. In appendix A, the CFT realization of $O(0)$, or more accurately, $O^+(0)$, is given. It is shown that the charge $Q_{0 \dots 0 a_1 \dots a_{2k-1}, b_1 \dots b_{s-1}}$ corresponding to (2.3) commutes with $O(0)$. So $a[E]$ constructed here is the same as (2.3).

The metric on the coset space $M = G[\text{ho}(1|2 : [3, 2])]/E$ is defined in group theory. Alternatively, we can use the operator $O(z)$ to get the same result. There is a one-to-one correspondence between $T_z(M) = \{v^M \partial_M | M = 1, \dots, \dim M\}$ and $K(z)$. For the given ∂_M , $\exists k_M(z) \in K(z)$ satisfying

$$\partial_M O(z) = i[k_M(z), O(z)]. \quad (2.9)$$

$\{k_M(z)\}$ compose the basis for $K(z)$, from which, one can define a special set of the coordinate on M

$$O(z) = e^{ik_M(0)z^M} O(0) e^{-ik_M(0)z^M}. \quad (2.10)$$

The metric on $T_z(M)$ can be induced from $K(z)$, i.e.

$$g_{MN}(z) = \langle k_M(z) | k_N(z) \rangle, \quad (2.11)$$

where $\langle k_M(z) | k_N(z) \rangle$ is the killing form. g_{MN} is $G[\text{ho}(1|2 : [3, 2])]$ invariant. Under the $G[\text{ho}(1|2 : [3, 2])]$ transformation,

$$O(z) \rightarrow gO(z)g^{-1} = O(z'). \quad (2.12)$$

$G[\text{ho}(1|2 : [3, 2])]$ generates the isometric transformation $z \rightarrow z'$ on M .

The tangent space on the coset space M is $\{t_A\}$. The group manifold of $\text{ho}(1|2 : [3, 2])$ is the manifold \mathbf{M} with the tangent space $\{t_\alpha\}$, $\dim \mathbf{M} = \dim \text{ho}(1|2 : [3, 2])$. The coordinate on \mathbf{M} is $Z_{\bar{M}}$, $ik_{\bar{M}}(Z) = \partial_{\bar{M}}g(Z)g(Z)^{-1}$, $G_{\bar{M}\bar{N}}(Z) = \langle k_{\bar{M}}(Z)|k_{\bar{N}}(Z) \rangle$.

$$\partial_{\bar{M}}O(Z) = i[k_{\bar{M}}(Z), O(Z)]. \quad (2.13)$$

When $k_{\bar{M}}(Z) \in E(Z)$, $\partial_{\bar{M}}O(Z) = 0$. Let $\{t_\alpha\}$ be a set of the orthogonal normalized basis of $\text{ho}(1|2 : [3, 2])$, one may assume $k_\alpha(Z) = g(Z)t_\alpha g(Z)^{-1}$. $k_{\bar{M}}(Z) = W_{\bar{M}}^\alpha(Z)k_\alpha(Z)$ and $k_\alpha(Z) = W_{\bar{\alpha}}^{\bar{M}}(Z)k_{\bar{M}}(Z)$ gives the vielbein on \mathbf{M} :

$$W_{\bar{M}}^\alpha W_{\bar{\beta}}^{\bar{M}} = \delta_{\bar{\beta}}^\alpha, \quad W_{\bar{M}}^\alpha W_{\bar{\alpha}}^{\bar{N}} = \delta_{\bar{\alpha}}^{\bar{N}}, \quad \eta_{\alpha\beta} W_{\bar{M}}^\alpha W_{\bar{N}}^\beta = G_{\bar{M}\bar{N}}. \quad (2.14)$$

$\eta_{\alpha\beta} = f_{\alpha\sigma}^\rho f_{\beta\rho}^\sigma = \langle t_\alpha|t_\beta \rangle$ is the killing metric for $\text{ho}(1|2 : [3, 2])$ with the suitable normalization assumed.³ Suppose $\partial_{\bar{N}}k_{\bar{M}}(Z) = \Gamma_{\bar{N}\bar{M}}^{\bar{L}}k_{\bar{L}}(Z)$, $\partial_{\bar{N}}k_\alpha(Z) = \phi_{\bar{N}\alpha}^\beta k_\beta(Z)$, there will be

$$\partial_{\bar{N}}W_{\bar{\alpha}}^{\bar{M}} + \Gamma_{\bar{N}\bar{L}}^{\bar{M}}W_{\bar{\alpha}}^{\bar{L}} - \phi_{\bar{N}\alpha}^\beta W_{\bar{\beta}}^{\bar{M}} = 0. \quad (2.15)$$

With the covariant derivative defined as $\mathcal{D}_{\bar{M}} = \partial_{\bar{M}} - \Gamma_{\bar{M}}$ and $\mathcal{D}_\alpha = W_{\bar{\alpha}}^{\bar{M}}(\partial_{\bar{M}} - \phi_{\bar{M}}) = \partial_\alpha - \phi_\alpha$, we have

$$\mathcal{D}_{\bar{M}_n} \cdots \mathcal{D}_{\bar{M}_2} \mathcal{D}_{\bar{M}_1} O(Z) = i^n [k_{\bar{M}_1}(Z), [k_{\bar{M}_2}(Z), \cdots [k_{\bar{M}_n}(Z), O(Z)] \cdots]], \quad (2.16)$$

$$\mathcal{D}_{\alpha_n} \cdots \mathcal{D}_{\alpha_2} \mathcal{D}_{\alpha_1} O(Z) = i^n [k_{\alpha_1}(Z), [k_{\alpha_2}(Z), \cdots [k_{\alpha_n}(Z), O(Z)] \cdots]]. \quad (2.17)$$

As is shown in appendix A, for $[Q_{0 \cdots 0a_1 \cdots a_s, b_1 \cdots b_{s+k}}, O(0)]$ with $k = 1, 3, \cdots$

$$\begin{aligned} & [Q_{0 \cdots 0a_1 \cdots a_s, b_1 \cdots b_{s+k}}, O(0)] \\ &= \sum_{r=0,2,\cdots,s}^{t=1,2,\cdots,2s+k-2r} g_{0 \cdots 0a_1 \cdots a_s, b_1 \cdots b_{s+k}}^{c_1 \cdots c_{2r+t}} [Q_{0c_1 \cdots c_r, c_{r+1} \cdots c_{2r+1}}, \cdots [Q_{0, c_{2r+t-1}}, [Q_{0, c_{2r+t}}, O(0)]] \cdots]. \end{aligned} \quad (2.18)$$

At the point Z , $O(Z) = g(Z)O(0)g(Z)^{-1}$, $Q_A(Z) = g(Z)Q_Ag(Z)^{-1}$,

$$\begin{aligned} & [Q_{0 \cdots 0a_1 \cdots a_s, b_1 \cdots b_{s+k}}(Z), O(Z)] \\ &= \sum_{r=0,2,\cdots,s}^{t=1,2,\cdots,2s+k-2r} g_{0 \cdots 0a_1 \cdots a_s, b_1 \cdots b_{s+k}}^{c_1 \cdots c_{2r+t}} [Q_{0c_1 \cdots c_r, c_{r+1} \cdots c_{2r+1}}(Z), \cdots [Q_{0, c_{2r+t-1}}(Z), [Q_{0, c_{2r+t}}(Z), O(Z)]] \cdots]. \end{aligned} \quad (2.19)$$

Since

$$\begin{aligned} & \mathcal{D}_{0, b_{s+k}} \mathcal{D}_{0, b_{s+k-1}} \cdots \mathcal{D}_{0a_1 \cdots a_s, b_1 \cdots b_{s+1}} O(Z) \\ &= i^k [Q_{0a_1 \cdots a_s, b_1 \cdots b_{s+1}}(Z), \cdots [Q_{0, b_{s+k-1}}(Z), [Q_{0, b_{s+k}}(Z), O(Z)]] \cdots], \end{aligned} \quad (2.20)$$

³Notice that the killing metric $\eta_{\alpha\beta}$ is indefinite having one sign for compact directions and the opposite for non-compact directions. $G(\text{ho}(1|2 : [3, 2]))$ is obviously not a compact group as one can see from its subgroup $\text{SO}(3, 2)$.

there will be

$$\begin{aligned} & \partial_{0 \dots 0 a_1 \dots a_s, b_1 \dots b_{s+k}} O(Z) \\ &= \sum_{r=0,2,\dots,s} \sum_{t=1,2,\dots,2s+k-2r} i^{1-t} g_{0 \dots 0 a_1 \dots a_s, b_1 \dots b_{s+k}}^{c_1 \dots c_{2r+t}} \mathcal{D}_{0, c_{2r+t}} \mathcal{D}_{0, c_{2r+t-1}} \dots \mathcal{D}_{0 c_1 \dots c_r, c_{r+1} \dots c_{2r+1}} O(Z). \end{aligned} \quad (2.21)$$

According to the definition, $\phi_{\bar{M}\gamma}^\beta$ and $W_\alpha^{\bar{M}}$ are invariant under the global higher spin transformation, so is their contraction ϕ_α . ϕ_α is a scalar, so it must be a constant on \mathbf{M} . (2.21) can be further rewritten as

$$\begin{aligned} & \partial_{0 \dots 0 a_1 \dots a_s, b_1 \dots b_{s+k}} O(Z) \\ &= \sum_{r=0,2,\dots,s} \sum_{t=1,2,\dots,2s+k-2r} G_{0 \dots 0 a_1 \dots a_s, b_1 \dots b_{s+k}}^{c_1 \dots c_{2r+t}} \partial_{0, c_{2r+t}} \partial_{0, c_{2r+t-1}} \dots \partial_{0 c_1 \dots c_r, c_{r+1} \dots c_{2r+1}} O(Z) \end{aligned} \quad (2.22)$$

for some constant $G_{0 \dots 0 a_1 \dots a_s, b_1 \dots b_{s+k}}^{c_1 \dots c_{2r+t}}$.

Just as the chiral constraint relates $\partial_{\bar{\theta}}$ with ∂_μ , here, $\partial_{0 \dots 0 a_1 \dots a_s, b_1 \dots b_{s+k}}$ is determined by $\partial_{0, c_{2r+t}} \partial_{0, c_{2r+t-1}} \dots \partial_{0 c_1 \dots c_r, c_{r+1} \dots c_{2r+1}}$. This is because $[Q_{A_1 \dots A_{s-1}, B_1 \dots B_{s-1}}(Z), O(Z)]|0\rangle$ are all in the 1-particle Hilbert space of the higher spin theory, for which

$$\begin{aligned} & \{[Q_{0, b_{s+k}}(Z), \dots [Q_{0, b_{s+2}}(Z), [Q_{0 a_1 \dots a_s, b_1 \dots b_{s+1}}(Z), O(Z)]] \dots]\} \\ & \sim \{[Q_{0 a_1 \dots a_s, b_1 \dots b_{s+1}}(Z), [Q_{0, b_{s+2}}(Z), \dots, [Q_{0, b_{s+k}}(Z), O(Z)]] \dots]\} \end{aligned} \quad (2.23)$$

compose the complete basis.⁴ In [12], by considering the zeroth level of the unfolded equation for the 0-form Φ in M , the similar result is also obtained. $\Phi = \Phi^{[a(0), b(0)]}$ is the lowest component of $\Phi^{[a(s+t), b(s)]}$. Generically, one may expect

$$\begin{aligned} & [Q_{0 \dots 0 a_1^p \dots a_{s_p}^p, b_1^p \dots b_{s_p+k_p}^p}(Z), \dots [Q_{0 \dots 0 a_1^1 \dots a_{s_1}^1, b_1^1 \dots b_{s_1+k_1}^1}(Z), O(Z)] \dots] \\ & \sim \sum \alpha(a_1 \dots a_s, b_1 \dots b_{s+k}) [Q_{0 a_1 \dots a_s, b_1 \dots b_{s+1}}(Z), \dots [Q_{0, b_{s+k-1}}(Z), [Q_{0, b_{s+k}}(Z), O(Z)]] \dots], \end{aligned} \quad (2.24)$$

where $\alpha(a_1 \dots a_s, b_1 \dots b_{s+k})$ are constants to be determined.

$$\begin{aligned} & \partial_{0 \dots 0 a_1^1 \dots a_{s_1}^1, b_1^1 \dots b_{s_1+k_1}^1} \dots \partial_{0 \dots 0 a_1^p \dots a_{s_p}^p, b_1^p \dots b_{s_p+k_p}^p} O(Z) \\ & \sim \sum \Lambda(a_1 \dots a_s, b_1 \dots b_{s+k}) \partial_{0, b_{s+k}} \partial_{0, b_{s+k-1}} \dots \partial_{0 a_1 \dots a_s, b_1 \dots b_{s+1}} O(Z). \end{aligned} \quad (2.25)$$

(2.25) is the $G[\text{ho}(1|2 : [3, 2])]$ -invariant differential operators on M , which will be useful in section 3.6 when we try to construct the theory with the global higher spin symmetry.

3 Theory with the local higher spin symmetry

In section 2, the background in \mathbf{M} is fixed to be the intrinsic geometry with $dW_0^\alpha - \frac{1}{2} f_{\beta\gamma}^\alpha W_0^\beta \wedge W_0^\gamma = 0$, which is invariant under the global higher spin transformation preserving W_0^α . To

⁴More precisely, it is $\{[Q_{0, b_{s+k}}(Z), \dots [Q_{0, b_{s+2}}(Z), O_{a_1 \dots a_s, b_1 \dots b_s}^s(Z)]] \dots]\}$ that forms the complete basis, but (2.23) is enough to generate $[Q_{A_1 \dots A_{s-1}, B_1 \dots B_{s-1}}(Z), O(Z)]$ since $[Q_Q(Z), O(Z)] = 0$ for t_Q in (2.3).

have the local higher spin symmetry, the 1-form W^α in \mathbf{M} should be dynamical. We will study the dynamics of the 1-form W^α and the 0-form H in \mathbf{M} . With the suitable rheonomy condition and the torsion constraint imposed, (W^α, H) in the whole \mathbf{M} is determined by $(W_\mu^{[a(s-1), b(0)]}, H)$ in AdS_4 . We then discuss the relation between the unfolded equation in group manifold approach and the unfolded equation in Vasiliev theory. We will also make a comment on theory with the global higher spin symmetry.

3.1 Higher spin theory on group manifold and the rheonomy condition

The 1-form W_M^α is the vielbein on \mathbf{M} . $W_M^\alpha W_\beta^{\bar{M}} = \delta_\beta^\alpha$, $W_M^\alpha W_\alpha^{\bar{N}} = \delta_M^{\bar{N}}$, $\eta_{\alpha\beta} W_M^\alpha W_N^\beta = G_{\bar{M}\bar{N}}$.⁵ The curvature 2-form is defined as

$$R^\alpha = dW^\alpha - \frac{1}{2} f_{\beta\gamma}^\alpha W^\beta \wedge W^\gamma. \quad (3.1)$$

It is convenient to use the 0-form $R_{\beta\gamma}^\alpha$ to parameterize $R_{\bar{M}\bar{N}}^\alpha$, $R_{\bar{M}\bar{N}}^\alpha = R_{\beta\gamma}^\alpha W_M^\beta W_N^\gamma$, $R_{\beta\gamma}^\alpha = R_{\bar{M}\bar{N}}^\alpha W_\beta^{\bar{M}} W_\gamma^{\bar{N}}$.

$$dW^\alpha = \frac{1}{2} (f_{\beta\gamma}^\alpha + R_{\beta\gamma}^\alpha) W^\beta \wedge W^\gamma = \frac{1}{2} \hat{f}_{\beta\gamma}^\alpha W^\beta \wedge W^\gamma, \quad (3.2)$$

where $\hat{f}_{\beta\gamma}^\alpha$ is the deformed structure constant. The Bianchi identity is

$$\partial_{[\gamma} \hat{f}_{\rho\sigma]}^\alpha + \hat{f}_{\beta[\gamma}^\alpha \hat{f}_{\rho\sigma]}^\beta = 0, \quad (3.3)$$

where $\partial_\gamma = W_\gamma^{\bar{M}} \partial_{\bar{M}}$. In addition, we can add the 0-form matter field H on \mathbf{M} ,

$$dH = H_\alpha W^\alpha \Leftrightarrow \partial_\alpha H = H_\alpha, \quad (3.4)$$

$$\partial_{[\rho} H_{\sigma]} + H_\alpha \hat{f}_{\rho\sigma}^\alpha = 0. \quad (3.5)$$

The group manifold \mathbf{M} is necessarily involved in the definition of $R_{\beta\gamma}^\alpha$ and H_α . (3.3) and (3.5) are defined in \mathbf{M} as well.

The definition (3.2) and (3.4) is invariant under the diffeomorphism transformation generated by $\xi^{\bar{M}}$,

$$\delta_\xi W_M^\alpha = \xi^{\bar{N}} \partial_{\bar{N}} W_M^\alpha + \partial_{\bar{M}} \xi^{\bar{N}} W_N^\alpha, \quad \delta_\xi \hat{f}_{\beta\gamma}^\alpha = \xi^{\bar{N}} \partial_{\bar{N}} \hat{f}_{\beta\gamma}^\alpha, \quad \delta_\xi H = \xi^{\bar{N}} \partial_{\bar{N}} H, \quad \delta_\xi H_\alpha = \xi^{\bar{N}} \partial_{\bar{N}} H_\alpha. \quad (3.6)$$

With

$$\epsilon^\alpha = \xi^{\bar{M}} W_M^\alpha, \quad \xi^{\bar{M}} = \epsilon^\alpha W_\alpha^{\bar{M}}, \quad (3.7)$$

(3.6) can be rewritten as

$$\delta_\epsilon W^\alpha = d\epsilon^\alpha + \hat{f}_{\beta\gamma}^\alpha \epsilon^\beta W^\gamma, \quad \delta_\epsilon \hat{f}_{\rho\sigma}^\alpha = \epsilon^\beta \partial_\beta \hat{f}_{\rho\sigma}^\alpha, \quad \delta_\epsilon H = \epsilon^\beta H_\beta, \quad \delta_\epsilon H_\alpha = \epsilon^\beta \partial_\beta H_\alpha, \quad (3.8)$$

which is the deformed local higher spin transformation.

$$\delta_{\epsilon_2} \delta_{\epsilon_1} - \delta_{\epsilon_1} \delta_{\epsilon_2} = \delta_{[\epsilon_2, \epsilon_1]}, \quad [\epsilon_2, \epsilon_1]^\alpha = \hat{f}_{\beta\gamma}^\alpha \epsilon_2^\gamma \epsilon_1^\beta. \quad (3.9)$$

The algebra is closed with the deformed structure constant $\hat{f}_{\beta\gamma}^\alpha$.

⁵Here, W_M^α is invertible, which is general enough to account for the 4d HS theory, in which, the relevant field is W_μ^a . Let $\{\alpha\} = \{\bar{\alpha}\} \cup \{a\}$, $\{\bar{M}\} = \{\bar{M}\} \cup \{\mu\}$, $W_\mu^a \sim e_\mu^a$ is usually required to be invertible, one can also suitably select $W_{\bar{M}}^{\bar{\alpha}}$ to make the whole $W_{\bar{M}}^\alpha$ invertible.

If for some Λ , $R_{\Lambda\gamma}^\alpha = 0$, $\hat{f}_{\Lambda\gamma}^\alpha = f_{\Lambda\gamma}^\alpha$, the local gauge transformation generated by ϵ^Λ is undeformed. It is necessary to require $R_{[a(1),b(1)]\gamma}^\alpha \equiv R_{(ab)\gamma}^\alpha = 0$ to make the local Lorentz transformation undeformed. Also, since H is a scalar, $H_{(ab)} = 0$ should hold so that $\delta_{\epsilon^{ab}}H = \epsilon^{ab}H_{(ab)} = 0$. From (3.3) and (3.5),

$$\delta_{\epsilon^{ab}}R_{\rho\sigma}^\alpha = \epsilon^{ab}\partial_{(ab)}R_{\rho\sigma}^\alpha = \epsilon^{ab}[f_{(ab)\beta}^\alpha R_{\rho\sigma}^\beta + f_{(ab)[\rho}^\beta R_{\sigma]\beta}^\alpha], \quad \delta_{\epsilon^{ab}}H_\alpha = \epsilon^{ab}\partial_{(ab)}H_\alpha = -\epsilon^{ab}f_{(ab)\alpha}^\beta H_\beta. \quad (3.10)$$

The evolution along the (ab) direction is a local Lorentz transformation, so the group manifold \mathbf{M} effectively reduces to the coset space $\mathcal{M} = G[\text{ho}(1|2 : [3, 2])]/\text{SO}(3, 1)$. Recall that in section 2, we have discussed the coset space $M = G[\text{ho}(1|2 : [3, 2])]/E$. For \mathbf{M} to reduce to M , there must be $R_{Q\gamma}^\alpha = 0$ so that the local gauge transformation generated by ϵ^Q is undeformed. However, at least in Vasiliev theory, $R_{(ab)\gamma}^\alpha = 0$ is valid but $R_{Q\gamma}^\alpha = 0$ does not necessarily hold.

When $\beta \neq (ab)$, $\partial_\beta \hat{f}_{\rho\sigma}^\alpha$ and $\partial_\beta H_\alpha$ cannot be uniquely determined by (3.3) and (3.5). Nevertheless, from (3.3) and (3.5), we have

$$\partial_\gamma R_{ab}^\alpha = \partial_{[b}R_{a]\gamma}^\alpha + \hat{f}_{\beta[\gamma}^\alpha \hat{f}_{ba]}^\beta \quad (3.11)$$

$$\partial_\gamma H_a = \partial_a H_\gamma + H_\alpha \hat{f}_{a\gamma}^\alpha \quad (3.12)$$

with $H_a = \partial_a H$. a represents the $[0, a]$ element of $\text{ho}(1|2 : [3, 2])$. Let

$$R_{ab;c_1 \dots c_n}^\alpha = \partial_{c_n} \dots \partial_{c_1} R_{ab}^\alpha, \quad H_{c_1 \dots c_n} = \partial_{c_n} \dots \partial_{c_1} H, \quad (3.13)$$

if

$$\begin{aligned} R_{\beta\gamma}^\alpha &= r_{\beta\gamma}^\alpha(R_{ab}^\sigma, R_{ab;c_1}^\sigma, \dots, H, H_{c_1}, \dots) \\ H_\gamma &= h_\gamma(R_{ab}^\sigma, R_{ab;c_1}^\sigma, \dots, H, H_{c_1}, \dots) \end{aligned} \quad (3.14)$$

with $r_{\beta\gamma}^\alpha$ and h_γ the polynomials of $R_{ab}^\sigma, R_{ab;c_1}^\sigma, \dots, H, H_{c_1}, \dots$ with the constant coefficients, then

$$\partial_\gamma R_{ab}^\alpha = \partial_{[b}R_{a]\gamma}^\alpha + \hat{f}_{\beta[\gamma}^\alpha \hat{f}_{ba]}^\beta = r_{ab;\gamma}^\alpha(R_{ab}^\sigma, R_{ab;c_1}^\sigma, \dots, H, H_{c_1}, \dots) \quad (3.15)$$

$$\partial_\gamma H_a = \partial_a H_\gamma + H_\alpha \hat{f}_{a\gamma}^\alpha = h_{a;\gamma}(R_{ab}^\sigma, R_{ab;c_1}^\sigma, \dots, H, H_{c_1}, \dots) \quad (3.16)$$

are also polynomials. Moreover, since

$$(\partial_\beta \partial_\gamma - \partial_\gamma \partial_\beta)F = \hat{f}_{\gamma\beta}^\alpha \partial_\alpha F, \quad (3.17)$$

$$(\partial_c \partial_\gamma - \partial_\gamma \partial_c)R_{ab}^\alpha = \hat{f}_{\gamma c}^\sigma \partial_\sigma R_{ab}^\alpha = \hat{f}_{\gamma c}^\sigma r_{ab;\sigma}^\alpha(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots), \quad (3.18)$$

$$(\partial_c \partial_\gamma - \partial_\gamma \partial_c)H_a = \hat{f}_{\gamma c}^\sigma \partial_\sigma H_a = \hat{f}_{\gamma c}^\sigma h_{a;\sigma}(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots), \quad (3.19)$$

so

$$\partial_\gamma R_{ab;c}^\alpha = \partial_\gamma \partial_c R_{ab}^\alpha = \partial_c r_{ab;\gamma}^\alpha - \hat{f}_{\gamma c}^\sigma r_{ab;\sigma}^\alpha = r_{ab;c\gamma}^\alpha \quad (3.20)$$

and

$$\partial_\gamma H_{ac} = \partial_\gamma H_{a;c} = \partial_\gamma \partial_c H_a = \partial_c h_{a;\gamma} - \hat{f}_{\gamma c}^\sigma h_{a;\sigma} = h_{a;c\gamma} = h_{ac\gamma} \quad (3.21)$$

are again polynomials. Subsequently, one can prove for $n = 0, 1, \dots$, we have

$$\begin{aligned}\partial_\gamma R_{ab;c_1 \dots c_n}^\alpha &= r_{ab;c_1 \dots c_n \gamma}^\alpha(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots), \\ \partial_\gamma H_{c_1 \dots c_n} &= h_{c_1 \dots c_n \gamma}(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots),\end{aligned}\quad (3.22)$$

or equivalently,

$$\begin{aligned}dR_{ab;c_1 \dots c_n}^\alpha &= r_{ab;c_1 \dots c_n \gamma}^\alpha(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots)W^\gamma, \\ dH_{c_1 \dots c_n} &= h_{c_1 \dots c_n \gamma}(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots)W^\gamma,\end{aligned}\quad (3.23)$$

where $r_{ab;c_1 \dots c_n \gamma}^\alpha$ and $h_{c_1 \dots c_n \gamma}$ are polynomials of $R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots$. Finally, we get the unfolded equation

$$\begin{aligned}dW^\alpha &= \frac{1}{2}(f_{\beta\gamma}^\alpha + r_{\beta\gamma}^\alpha)W^\beta \wedge W^\gamma, \\ dR_{ab;c_1 \dots c_n}^\alpha &= r_{ab;c_1 \dots c_n \gamma}^\alpha W^\gamma, \\ dH_{c_1 \dots c_n} &= h_{c_1 \dots c_n \gamma} W^\gamma,\end{aligned}\quad (3.24)$$

with $n = 0, 1, \dots$. $r_{\beta\gamma}^\alpha$, $r_{ab;c_1 \dots c_n \gamma}^\alpha$ and $h_{c_1 \dots c_n \gamma}$ are functions of $R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots$. From $(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots)$ at one point, (W^α, H) on the whole \mathbf{M} can be determined up to a gauge transformation. (3.24) is invariant under the local gauge transformation (3.8) which can now be explicitly written as

$$\begin{aligned}\delta_\epsilon W^\alpha &= d\epsilon^\alpha + \hat{f}_{\sigma\gamma}^\alpha \epsilon^\sigma W^\gamma, \\ \delta_\epsilon R_{ab;c_1 \dots c_n}^\alpha &= \epsilon^\sigma r_{ab;c_1 \dots c_n \sigma}^\alpha, \\ \delta_\epsilon H_{c_1 \dots c_n} &= \epsilon^\sigma h_{c_1 \dots c_n \sigma}.\end{aligned}\quad (3.25)$$

$(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots)$ forms a complete higher spin multiplet.

(3.14) is the rheonomy condition in higher spin theory. This is the most generic rheonomy condition requiring that the curvature $(R_{\beta\gamma}^\alpha, H_\gamma)$ is determined by its inner components (R_{ab}^α, H_a) as well as their inner derivatives. The condition, together with the Bianchi identity, gives the unfolded equation. The rheonomy condition in supergravity (1.1) is a special situation, in which, the dependence on the inner derivatives vanishes. Therefore, supergravity does not contain the higher derivative interactions.

The parameterization (3.14) should satisfy the Bianchi identity

$$\partial_{[\gamma} R_{\rho\sigma]}^\alpha + \hat{f}_{\beta[\gamma}^\alpha \hat{f}_{\rho\sigma]}^\beta = 0, \quad \partial_{[\gamma} H_{\beta]} + H_\alpha \hat{f}_{\gamma\beta}^\alpha = 0. \quad (3.26)$$

With (3.22) and (3.14) plugged in (3.26), we get

$$F_{[\gamma\rho\sigma]}^\alpha(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots) = 0, \quad F_{[\beta\gamma]}(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots) = 0, \quad (3.27)$$

where $F_{[\gamma\rho\sigma]}^\alpha$ and $F_{[\beta\gamma]}$ are also polynomials of $R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots$. (3.27) gives the 4d equations of motion for $(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots)$. For the randomly selected function $(r_{\beta\gamma}^\alpha, h_\gamma)$, (3.27) only has the trivial solution $R_{ab}^\beta = H = 0$. $(r_{\beta\gamma}^\alpha, h_\gamma)$ should be chosen to

allow as many on-shell degrees of freedom as possible. In this sense, (3.27) determines both $(r_{\beta\gamma}^\alpha, h_\gamma)$ and the $4d$ equations of motion.

To guarantee the local Lorentz invariance, in (3.14), $R_{(ab)\beta}^\alpha = H_{(ab)} = 0$. Since

$$\begin{aligned}\partial_{(ab)}R_{\rho\sigma}^\alpha &= \frac{\partial R_{\rho\sigma}^\alpha}{\partial R_{de;c_1\cdots c_n}^\beta}\partial_{(ab)}R_{de;c_1\cdots c_n}^\beta + \frac{\partial R_{\rho\sigma}^\alpha}{\partial H_{c_1\cdots c_n}}\partial_{(ab)}H_{c_1\cdots c_n}, \\ \partial_{(ab)}H_\alpha &= \frac{\partial H_\alpha}{\partial R_{de;c_1\cdots c_n}^\beta}\partial_{(ab)}R_{de;c_1\cdots c_n}^\beta + \frac{\partial H_\alpha}{\partial H_{c_1\cdots c_n}}\partial_{(ab)}H_{c_1\cdots c_n},\end{aligned}\quad (3.28)$$

where $\partial_{(ab)}R_{\rho\sigma}^\alpha$, $\partial_{(ab)}H_\alpha$, $\partial_{(ab)}R_{de;c_1\cdots c_n}^\beta$ and $\partial_{(ab)}H_{c_1\cdots c_n}$ are all standard local Lorentz transformations, the coefficients in $r_{\rho\sigma}^\alpha$ and h_α should be the Lorentz invariants. In fact, (3.28) are also included in (3.26), so the Lorentz invariance of $r_{\rho\sigma}^\alpha$ and h_α is also the requirement of the Bianchi identity if $R_{(ab)\beta}^\alpha = H_{(ab)} = 0$.

We only considered the equation (3.24) on group manifold \mathbf{M} , since in that space, the diffeomorphism transformation and the local gauge transformation are in one-to-one correspondence. As the universal property of the unfolded equation [15], (3.24) is well-defined in space m with $\dim m \geq 4$. If $\dim m > \dim \mathbf{M}$, different diffeomorphism transformations may be realized as the same gauge transformation, i.e. there are flat directions with $\xi^{\bar{M}}W_{\bar{M}}^\alpha = 0$; if $\dim m < \dim \mathbf{M}$, some gauge transformation does not have the diffeomorphism realization like that in AdS_4 .

The initial value is $(R_{ab}^\alpha, R_{ab;c_1}^\alpha, \dots, H, H_{c_1}, \dots)$ at one point, it is desirable to express it in terms of (W_μ^α, H) as well as its $4d$ derivatives at that point.

$$\begin{aligned}R_{\mu\nu}^\alpha &= r_{\beta\gamma}^\alpha W_\mu^\beta W_\nu^\gamma \\ \partial_\lambda R_{\mu\nu}^\alpha &= \left(\frac{\partial r_{\beta\gamma}^\alpha}{\partial R_{ab;c_1\cdots c_n}^\sigma} r_{ab;c_1\cdots c_n}^\sigma + \frac{\partial r_{\beta\gamma}^\alpha}{\partial H_{c_1\cdots c_n}} h_{c_1\cdots c_n} \right) W_\lambda^\rho W_\mu^\beta W_\nu^\gamma + r_{\beta\gamma}^\alpha \partial_\lambda W_{[\mu}^\beta W_{\nu]}^\gamma \\ &\dots \\ H_\mu &= h_\alpha W_\mu^\alpha \\ \partial_\lambda H_\mu &= \left(\frac{\partial h_\alpha}{\partial R_{ab;c_1\cdots c_n}^\sigma} r_{ab;c_1\cdots c_n}^\sigma + \frac{\partial h_\alpha}{\partial H_{c_1\cdots c_n}} h_{c_1\cdots c_n} \right) W_\lambda^\rho W_\mu^\alpha + h_\alpha \partial_\lambda W_\mu^\alpha \\ &\dots\end{aligned}\quad (3.29)$$

r and h are functions of $(R_{ab;c_1\cdots c_n}^\alpha, H_{c_1\cdots c_n})$. In (3.29), the unknowns are $(R_{ab;c_1\cdots c_n}^\alpha, H_{c_1\cdots c_n})$, while the number of equations is the same as the number of the degrees of freedom of $(R_{\mu\nu;\lambda_1\cdots\lambda_n}^\alpha, H_{\lambda_1\cdots\lambda_n})$, where $\mu, \nu = 1, 2, 3, 4$. (3.11) and (3.12) also impose constraints on the off-shell $(R_{ab;c_1\cdots c_n}^\alpha, H_{c_1\cdots c_n})$ to make it have the same number of degrees of freedom as $(R_{\mu\nu;\lambda_1\cdots\lambda_n}^\alpha, H_{\lambda_1\cdots\lambda_n})$, so in principle, from (3.29), $(R_{ab;c_1\cdots c_n}^\alpha, H_{c_1\cdots c_n})$ can be solved in terms of $(W_\mu^\alpha, \partial_{\nu_1} W_\mu^\alpha, \dots, H, \partial_{\nu_1} H, \dots)$.

$$\begin{aligned}R_{ab;c_1\cdots c_n}^\alpha &= g_{ab;c_1\cdots c_n}^\alpha(W_\mu^\sigma, \partial_{\nu_1} W_\mu^\sigma, \dots, H, \partial_{\nu_1} H, \dots), \\ H_{c_1\cdots c_n} &= q_{c_1\cdots c_n}(W_\mu^\sigma, \partial_{\nu_1} W_\mu^\sigma, \dots, H, \partial_{\nu_1} H, \dots).\end{aligned}\quad (3.30)$$

The local gauge transformation of (W_μ^α, H) in AdS_4 is

$$\begin{aligned}\delta_\epsilon W_\mu^\alpha &= \partial_\mu \epsilon^\alpha + \hat{f}_{\sigma\gamma}^\alpha(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots) \epsilon^\sigma W_\mu^\gamma, \\ \delta_\epsilon H &= \epsilon^\sigma h_\sigma(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots).\end{aligned}\quad (3.31)$$

With (3.30) plugged in (3.31),

$$\begin{aligned}\delta_\epsilon W_\mu^\alpha &= \partial_\mu \epsilon^\alpha + u_{\sigma\gamma}^\alpha(W_\mu^\sigma, \partial_{\nu_1} W_\mu^\sigma, \dots, H, \partial_{\nu_1} H, \dots) \epsilon^\sigma W_\mu^\gamma, \\ \delta_\epsilon H &= \epsilon^\sigma v_\sigma(W_\mu^\sigma, \partial_{\nu_1} W_\mu^\sigma, \dots, H, \partial_{\nu_1} H, \dots)\end{aligned}\quad (3.32)$$

gives the local gauge transformation rule of the matter-gravity coupled system (W_μ^α, H) in AdS_4 .

Since

$$(R_{ab}^\beta, R_{ab;c_1}^\beta, \dots, H, H_{c_1}, \dots) \sim (W_\mu^\sigma, \partial_{\nu_1} W_\mu^\sigma, \dots, H, \partial_{\nu_1} H, \dots), \quad (3.33)$$

(W^α, H) on the whole \mathbf{M} is determined by the on-shell $(W_\mu^\sigma, \partial_{\nu_1} W_\mu^\sigma, \dots, H, \partial_{\nu_1} H, \dots)$ at one point, or equivalently, the on-shell (W_μ^α, H) in AdS_4 . This is the rheonomy in higher spin theory. As is shown in section 2, although the space is \mathbf{M} with the infinite dimension, the physical Hilbert space is still the same as the $4d$ higher spin theory. Imposing the rheonomy condition is a way to project out the physical degrees of freedom.

3.2 Group manifold approach to supergravity

In this subsection, we will give a review of the group manifold approach for supergravity [1–4]. Some modification is made so that supergravity is treated in the same way as the above discussed higher spin theory.

For $\mathcal{N} = 1$ supergravity in $R^{3,1}$, the coordinate in group manifold is $(x^\mu, x^{\mu\nu}, \theta^\chi)$, the associated 1-form is $\nu^A = (\omega^{ab}, e^a, \psi^\alpha)$,⁶ and the 0-form matter field is H . We have

$$d\nu^A = \frac{1}{2} \hat{f}_{BC}^A \nu^B \wedge \nu^C, \quad dH = H_A \nu^A, \quad (3.34)$$

$$\partial_{[E} \hat{f}_{BC]}^A + \hat{f}_{D[E}^A \hat{f}_{BC]}^D = 0, \quad \partial_{[A} H_{B]} + H_C \hat{f}_{AB}^C = 0, \quad (3.35)$$

where $\hat{f}_{BC}^A = f_{BC}^A + R_{BC}^A$, $H_B = (H_a, H_{(ab)}, H_\alpha)$. f_{BC}^A is the structure constant of the super Poincaré group $\text{Osp}(4|1)$. (3.34) is invariant under the diffeomorphism transformation in group manifold generated by $\xi^{\bar{M}} = (\xi^{\mu\nu}, \xi^\mu, \xi^\chi)$

$$\delta_\xi \nu_M^A = \xi^{\bar{N}} \partial_{\bar{N}} \nu_M^A + \partial_{\bar{M}} \xi^{\bar{N}} \nu_N^A, \quad \delta_\xi \hat{f}_{BC}^A = \xi^{\bar{N}} \partial_{\bar{N}} \hat{f}_{BC}^A, \quad \delta_\xi H = \xi^{\bar{N}} \partial_{\bar{N}} H, \quad \delta_\xi H_A = \xi^{\bar{N}} \partial_{\bar{N}} H_A, \quad (3.36)$$

which, when written in terms of the components, are local Lorentz transformation, the $4d$ diffeomorphism transformation and the supersymmetry transformation respectively. With $\epsilon^A = \xi^{\bar{M}} \nu_M^A$, (3.36) can be rewritten as

$$\delta_\epsilon \nu^A = d\epsilon^A + \hat{f}_{BC}^A \epsilon^B \nu^C, \quad \delta_\epsilon \hat{f}_{BC}^A = \epsilon^D \partial_D \hat{f}_{BC}^A, \quad \delta_\epsilon H = \epsilon^D H_D, \quad \delta_\epsilon H_A = \epsilon^D \partial_D H_A. \quad (3.37)$$

⁶Here, α is the spinor index and should be distinguished from α in the rest sections, which represents the adjoint representation of $\text{ho}(1|2 : [3, 2])$. Also, α here is equivalent to the spinor index $(\alpha, \dot{\alpha})$ in section 3.

Until now, no dynamics is involved at all. The dynamical information is brought by imposing the suitable constraints on R_{BC}^A and H_A . Here, the constraints that will be imposed are

(a) Factorization condition $R_{(ab)C}^A = 0 = H_{(ab)}$;

(b) Rheonomy condition and the torsion constraint:

$$(i) \quad R_{BC}^A = r_{BC}^A(R_{ab}^{cd}, R_{ab,c_1}^{cd}, \dots, R_{ab}^\alpha, R_{ab,c_1}^\alpha, \dots, H, H_{c_1}, \dots, H_\alpha, H_{\alpha;c_1}, \dots), \quad (3.38)$$

or

$$(ii) \quad \begin{aligned} R_{BC}^A &= r_{BC}^A(R_{ab}^{cd}, R_{ab,c_1}^{cd}, \dots, R_{ab}^\alpha, R_{ab,c_1}^\alpha, \dots, H, H_{c_1}, \dots) \\ H_A &= h_A(R_{ab}^{cd}, R_{ab,c_1}^{cd}, \dots, R_{ab}^\alpha, R_{ab,c_1}^\alpha, \dots, H, H_{c_1}, \dots). \end{aligned} \quad (3.39)$$

(a) is imposed so that the local Lorentz transformation is undeformed. In (b), the rheonomy condition requires that the lower index of the independent fields can only contain a so that the whole dynamics in group manifold is determined by that in a $4d$ submanifold; torsion constraint requires that the upper index cannot be a so that ω^{ab} can be solved in terms of the rest fields. There are two possibilities. In (i), the final dynamical fields are $(e_\mu^a, \psi_\mu^\alpha, H, H_\alpha)$ in M_4 , which is the situation for $\mathcal{N} = 1$ supergravity coupled to the WZ matter. In (ii), the dynamical fields are $(e_\mu^a, \psi_\mu^\alpha, H)$ in M_4 like that in higher spin theory.

r_{BC}^A and h_A are polynomials, the coefficients of which should be selected so that some scaling relation is respected [3]. The weight of t^A is denoted as $w(A)$, $w(a) = 1$, $w(ab) = 0$, $w(\alpha) = 1/2$. The super Poincaré algebra $[t_{A_1}, t_{A_2}] = i f_{A_1 A_2}^{A_3} t_{A_3}$ is invariant under

$$t_{A_i} \rightarrow v^{-w(A_i)} t_{A_i} \quad (3.40)$$

The 0-forms H_A and R_{BC}^A have the weight $-w(A)$ and $w(A) - w(B) - w(C)$ as follows

$$\begin{array}{cccccccccc} H_a & H_\alpha & R_{ab}^{cd} & R_{a\alpha}^{cd} & R_{\alpha\beta}^{cd} & R_{ab}^c & R_{a\alpha}^c & R_{\alpha\beta}^c & R_{ab}^\gamma & R_{a\alpha}^\gamma & R_{\alpha\beta}^\gamma \\ -1 & -\frac{1}{2} & -2 & -\frac{3}{2} & -1 & -1 & -\frac{1}{2} & 0 & -\frac{3}{2} & -1 & -\frac{1}{2} \end{array}$$

Especially, $(R_{ab}^{cd}, R_{ab}^\alpha, H, H_a, H_\alpha)$ have the weight $(-2, -3/2, 0, -1, -1/2)$. (ii) cannot satisfy the scaling relation thus should be ruled out. For (i), with the H_A odd terms dropped, the most general form of r_{BC}^A is

$$\begin{aligned} R_{a\alpha}^{bc} &= r_{a\alpha}^{bc} |_{\beta}^{de} R_{de}^{\beta} + r_{a\alpha}^{bc} |^{\beta, d} H_{\beta} H_d, \\ R_{ab}^c &= r_{ab}^c |^{\alpha, \beta} H_{\alpha} H_{\beta}, \\ R_{a\alpha}^{\beta} &= r_{a\alpha}^{\beta} |^{\rho, \sigma} H_{\rho} H_{\sigma}, \\ R_{\alpha\beta}^{cd} &= r_{\alpha\beta}^{cd} |^{\rho, \sigma} H_{\rho} H_{\sigma}, \\ R_{a\alpha}^c &= R_{\alpha\beta}^{\gamma} = R_{\alpha\beta}^c = 0, \end{aligned} \quad (3.41)$$

where $r_{**}^{*|**} = r_{**}^{*|**}(H)$ are functions of H since H has the weight 0. $r_{**}^{*|**}$ should be a Lorentz invariant to preserve the local Lorentz invariance. Although the torsion

constraint is also imposed, R_{AB}^a does not need to vanish, see for example [16]. However, if $H_\alpha = H_a = 0$, $R_{AB}^a = 0$, so in pure supergravity case, we do have $R_{AB}^a = 0$. Due to the scaling relation, the rheonomy condition is greatly simplified. For supergravity in AdS_4 with the symmetry group $\text{Osp}(4|1)$, a constant L with the weight 1 is involved. $L \rightarrow \infty$ gives the flat space limit, so only the L^{-n} terms with $n \geq 0$ are allowed in rheonomy condition. (3.41) remains valid.

(3.41) should satisfy the Bianchi identity

$$\partial_{[E} R_{BC]}^A + f_{D[E}^A R_{BC]}^D + R_{D[E}^A f_{BC]}^D + R_{D[E}^A R_{BC]}^D = 0, \quad \partial_{[A} H_{B]} + H_C f_{AB}^C + H_C R_{AB}^C = 0. \quad (3.42)$$

In pure supergravity situation with $H = 0$, r_{BC}^A becomes

$$\begin{aligned} R_{a\alpha}^{bc} &= r_{a\alpha}^{bc|de} R_{de}^\beta, \\ R_{ab}^c &= R_{a\alpha}^\beta = R_{\alpha\beta}^{cd} = R_{a\alpha}^c = R_{\alpha\beta}^\gamma = R_{\alpha\beta}^c = 0. \end{aligned} \quad (3.43)$$

(3.42) reduces to

$$\partial_{(ab)} R_{BC}^A = f_{(ab)D}^A R_{BC}^D - f_{(ab)C}^D R_{BD}^A - f_{(ab)B}^D R_{DC}^A, \quad (3.44)$$

$$f_{(ef)[b}^a R_{cd]}^{ef} = 0, \quad \partial_{[a} R_{bc]}^\alpha = 0, \quad \partial_{[c} R_{de]}^{ab} + R_{\alpha[c}^{ab} R_{de]}^\alpha = 0, \quad (3.45)$$

$$\partial_\beta R_{bc}^\alpha + f_{(ad)\beta}^\alpha R_{bc}^{(ad)} = 0, \quad \partial_\alpha R_{cd}^{ab} + \partial_{[c} R_{d]\alpha}^{ab} = 0, \quad (3.46)$$

$$f_{\beta\alpha}^a R_{bc}^\beta + f_{(ef)[b}^a R_{c]\alpha}^{ef} = 0, \quad R_{ac}^\alpha f_{\beta\gamma}^\alpha + f_{(ab)[\beta}^\alpha R_{\gamma]c}^{ab} = 0, \quad \partial_{[\alpha} R_{\beta]c}^{ab} + R_{dc}^{ab} f_{\alpha\beta}^d = 0. \quad (3.47)$$

(3.44) gives the Lorentz transformation of R_{BC}^A , which can be preserved in $r_{BC}^A(R_{bc}^{ad}, R_{bc}^\beta)$ if $r_{a\alpha}^{bc|de}$ is a Lorentz scalar. (3.45) are Bianchi identities in $4d$. (3.46) gives the evolution of $(R_{bc}^{ad}, R_{bc}^\beta)$ along the α direction. With (3.46) plugged in (3.47), $r_{a\alpha}^{bc|de}$ can be fixed and the $4d$ equations of motion

$$R_{ab}^{cb} - \frac{1}{2} \delta_a^c R_{db}^{db} = 0, \quad \varepsilon^{abcd} (\gamma_5 \gamma_b)^\alpha_\beta R_{cd}^\beta = 0 \quad (3.48)$$

come out. If we use the on-shell \tilde{R}_{ab}^{cb} and \tilde{R}_{cd}^β satisfying (3.48) to parameterize r_{BC}^A , (3.47) will hold automatically. This is in analogy with Vasiliev theory, with $R_{\beta\gamma}^\alpha$ parametrized by the 0-form $\Phi^{\tilde{\alpha}}$ in the twisted-adjoint representation of the higher spin algebra, the Bianchi identity is satisfied for the arbitrary $\Phi^{\tilde{\alpha}}$.

Written as the unfolded equation,

$$\begin{aligned} d\nu^A &= \frac{1}{2} (f_{BC}^A + r_{BC}^A) \nu^B \wedge \nu^C, \\ dR_{ab;c_1 \dots c_n}^{cd} &= r_{ab;c_1 \dots c_n A}^{cd} \nu^A, & dR_{ab;c_1 \dots c_n}^\alpha &= r_{ab;c_1 \dots c_n A}^\alpha \nu^A, \\ dH_{c_1 \dots c_n} &= h_{c_1 \dots c_n A} \nu^A, & dH_{\alpha;c_1 \dots c_n} &= h_{\alpha;c_1 \dots c_n A} \nu^A, \end{aligned} \quad (3.49)$$

where r, h are all determined by $r_{BC}^A = r_{BC}^A(R_{ab}^{cd}, R_{ab}^\alpha, H_a, H_\alpha)$ and are functions of $(R_{ab;c_1 \dots c_n}^{cd}, R_{ab;c_1 \dots c_n}^\alpha, H_{c_1 \dots c_n}, H_{\alpha;c_1 \dots c_n})$. With the on-shell $(R_{ab;c_1 \dots c_n}^{cd}, R_{ab;c_1 \dots c_n}^\alpha, H_{c_1 \dots c_n}, H_{\alpha;c_1 \dots c_n})$,

$H_{\alpha;c_1 \dots c_n}$) given at one point, $(\nu^A, R_{ab;c_1 \dots c_n}^{cd}, R_{ab;c_1 \dots c_n}^\alpha, H_{c_1 \dots c_n}, H_{\alpha;c_1 \dots c_n})$ on the whole \mathbf{M} can be solved. The local gauge transformation is

$$\begin{aligned} \delta_\epsilon \nu^A &= d\epsilon^A + \hat{f}_{BC}^A \epsilon^B \nu^C, \\ \delta_\epsilon R_{ab;c_1 \dots c_n}^{cd} &= \epsilon^A r_{ab;c_1 \dots c_n A}^{cd}, & \delta_\epsilon R_{ab;c_1 \dots c_n}^\alpha &= \epsilon^A r_{ab;c_1 \dots c_n A}^\alpha, \\ \delta_\epsilon H_{c_1 \dots c_n} &= \epsilon^A h_{c_1 \dots c_n A}, & \delta_\epsilon H_{\alpha;c_1 \dots c_n} &= \epsilon^A h_{\alpha;c_1 \dots c_n A}. \end{aligned} \quad (3.50)$$

$(R_{ab}^{cd}, R_{ab;c_1}^{cd}, \dots, R_{ab}^\alpha, R_{ab;c_1}^\alpha, \dots, H, H_{c_1}, \dots, H_\alpha, H_{\alpha;c_1}, \dots)$ compose the complete super-symmetry multiplet.

In addition to the 1-form ν^A , the 0-form multiplet is introduced, forming the representation of the deformed local super Poincaré transformation. The physical interpretation of the 0-form is the curvature and the matter field plus their derivatives. This is in the same spirit as the higher spin theory. Different from the higher spin theory, rheonomy condition (3.41) only contains $R_{ab}^{cd}, R_{ab}^\alpha, H_\alpha, H_\alpha$, so the infinite length 0-form multiplet does not enter into the $4d$ equations of motion. As a result, the equations of motion for $(e_\mu^a, \psi_\mu^\alpha, H, H_\alpha)$ do not contain the higher order derivatives. One may similarly make a robust requirement $R_{\beta\gamma}^\alpha = r_{\beta\gamma}^\alpha(R_{ab}^\sigma, H)$ and $H_\gamma = h_\gamma(R_{ab}^\sigma, H)$ in higher spin theory. However, such $(r_{\beta\gamma}^\alpha, h_\gamma)$ may only allow the trivial solution $R_{ab}^\sigma = H = 0$ when the Bianchi identity is imposed, no matter how coefficients in $(r_{\beta\gamma}^\alpha, h_\gamma)$ are adjusted.

Again,

$$\begin{aligned} &(R_{ab}^{cd}, R_{ab;c_1}^{cd}, \dots, R_{ab}^\alpha, R_{ab;c_1}^\alpha, \dots, H, H_{c_1}, \dots, H_\alpha, H_{\alpha;c_1}, \dots) \\ &\sim (e_\mu^b, \partial_{\nu_1} e_\mu^b, \dots, \psi_\mu^\alpha, \partial_{\nu_1} \psi_\mu^\alpha, \dots, H, \partial_{\nu_1} H, \dots, H_\alpha, \partial_{\nu_1} H_\alpha, \dots). \end{aligned} \quad (3.51)$$

With the on-shell $(e_\mu^a, \psi_\mu^\alpha, H, H_\alpha)$ given on M_4 , (ν^A, H) on the whole group manifold can be determined up to a gauge transformation.

The dynamics is entirely encoded in function $r_{BC}^A(R_{bc}^{ad}, R_{bc}^\beta, H, H_\alpha)$. By setting H to 0, we obtain the pure supergravity situation. Alternatively, one can consider the dynamics of the 0-form matter on the fixed supergravity background by setting r_{BC}^A to 0. With $\hat{f}_{BC}^A = f_{BC}^A$, (3.34) and (3.35) reduce to

$$d\nu_0^A = \frac{1}{2} f_{BC}^A \nu_0^B \wedge \nu_0^C, \quad dH = H_A \nu_0^A, \quad \partial_{[A} H_{B]} + H_C f_{AB}^C = 0. \quad (3.52)$$

ν_0^A describes the intrinsic geometry of the group manifold. The allowed gauge transformation parameter ϵ_0^A should make ν_0^A invariant

$$\delta_{\epsilon_0} \nu_0^A = \partial_{\bar{M}} \epsilon_0^A + f_{BC}^A \epsilon_0^B \nu_0^C = 0. \quad (3.53)$$

ϵ_0^A generates the global super Poincaré transformation on group manifold.

$$\delta_{\epsilon_0} H = \xi_0^{\bar{M}} \partial_{\bar{M}} H = \epsilon_0^D H_D, \quad \delta_{\epsilon_0} H_A = \xi_0^{\bar{M}} \partial_{\bar{M}} H_A = \epsilon_0^D \partial_D H_A. \quad (3.54)$$

$\xi_0^{\bar{M}} \nu_0^D = \epsilon_0^D$. Still, $H_{(ad)} = 0$.

$$\begin{aligned} \partial_{(ad)} H &= 0, \\ \partial_{(ad)} H_c + H_b f_{(ad)c}^b &= 0, \\ \partial_{(ad)} H_\alpha + H_\beta f_{(ad)\alpha}^\beta &= 0. \end{aligned} \quad (3.55)$$

Evolution along (ad) direction is a Lorentz transformation. One cannot assume H_α is the function of $(H, H_{c_1}, H_{c_1 c_2}, \dots)$, since the scaling relation is not respected. Let $\alpha = (\lambda, \dot{\lambda})$, one can at most require $H_{\dot{\lambda}} = 0$, which is the chiral constraint for superfield.

3.3 Imposing the torsion constraint in higher spin theory

Back to higher spin theory, a further reduction of (3.33) can be made by imposing the following torsion constraint

$$\begin{aligned} R_{\beta\gamma}^\alpha &= r_{\beta\gamma}^\alpha(R_{ab}^{[a(s-1),b(s-1)]}, R_{ab;c_1}^{[a(s-1),b(s-1)]}, \dots, H, H_{c_1}, \dots), \\ H_\gamma &= h_\gamma(R_{ab}^{[a(s-1),b(s-1)]}, R_{ab;c_1}^{[a(s-1),b(s-1)]}, \dots, H, H_{c_1}, \dots). \end{aligned} \quad (3.56)$$

Namely, in (3.14), σ is restricted to $[a(s-1), b(s-1)]$ with $s = 2, 4, \dots$. In (3.29), the number of equations is equal to the number of degrees of freedom of $(R_{ab, c_1 \dots c_n}^\alpha, H_{c_1 \dots c_n})$ but the number of unknowns is equal to the degrees of freedom of $(R_{ab, c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ now, so effectively, there will be some constraints imposed on (W_μ^α, H) in AdS_4 whose number is equal to the degrees of freedom of $R_{ab}^{[a(s-1), b(t)]}$ with $0 \leq t \leq s-2$. It is expected that by solving these constraints, $W_\mu^{[a(s-1), b(t+1)]}$ can be expressed in terms of $(W_\mu^{[a(s-1), b(0)]}, H)$. In fact, at least in free theory limit, imposing the torsion constraint $R_{ab}^{[a(s-1), b(t)]} = 0$ for $0 \leq t \leq s-2$ can indeed make $W_\mu^{[a(s-1), b(t+1)]}$ solved in terms of $W_\mu^{[a(s-1), b(0)]}$ [17]. (3.33) then reduces to

$$\begin{aligned} &(R_{ab}^{[a(s-1), b(s-1)]}, R_{ab; c_1}^{[a(s-1), b(s-1)]}, \dots, H, H_{c_1}, \dots) \\ &\sim (W_\mu^{[a(s-1), b(0)]}, \partial_{\nu_1} W_\mu^{[a(s-1), b(0)]}, \dots, H, \partial_{\nu_1} H, \dots). \end{aligned} \quad (3.57)$$

With $W_\mu^{[a(s-1), b(t+1)]}$ written in terms of $(W_\mu^{[a(s-1), b(0)]}, H)$, (3.32) becomes

$$\begin{aligned} \delta_\epsilon W_\mu^{[a(s-1), b(0)]} &= \partial_\mu \epsilon^{[a(s-1), b(0)]} + \epsilon^\sigma m_{\sigma\gamma}^{[a(s-1), b(0)]} (W_\mu^{[a(r-1), b(0)]}, \partial_{\nu_1} W_\mu^{[a(r-1), b(0)]}, \dots, \\ &H, \partial_{\nu_1} H, \dots) w_\mu^\gamma (W_\mu^{[a(r-1), b(0)]}, \partial_{\nu_1} W_\mu^{[a(r-1), b(0)]}, \dots, H, \partial_{\nu_1} H, \dots), \\ \delta_\epsilon H &= \epsilon^\sigma n_\sigma (W_\mu^{[a(r-1), b(0)]}, \partial_{\nu_1} W_\mu^{[a(r-1), b(0)]}, \dots, H, \partial_{\nu_1} H, \dots), \end{aligned} \quad (3.58)$$

which is the local gauge transformation rule of $(W_\mu^{[a(s-1), b(0)]}, H)$ in AdS_4 . In free theory limit, it is

$$h^{\mu_1 \dots \mu_s} = W_{a_1}^{\mu_1} W_{a_{s-1}}^{\mu_{s-1}} W_{a_s}^{\mu_s} g^{a_s a} W_a^\mu W_\mu^{a_1 \dots a_{s-1}, 0 \dots 0} \quad (3.59)$$

that will finally appear in equations of motion and the gauge transformation. One may expect in interacting case, the final dynamics is also expressed in terms of some $h^{\mu_1 \dots \mu_s}$, which can be a more complicated combination of $W_\mu^{a_1 \dots a_{s-1}, 0 \dots 0}$. The frame-like formulation reduces to the metric-like formulation.

Altogether, the complete equations are

$$\hat{f}_{\beta\gamma}^\alpha = \hat{f}_{\beta\gamma}^\alpha(R_{ab}^{[a(t-1), b(t-1)]}, R_{ab; c_1}^{[a(t-1), b(t-1)]}, \dots, H, H_{c_1}, \dots), \quad (3.60)$$

$$H_\gamma = h_\gamma(R_{ab}^{[a(t-1), b(t-1)]}, R_{ab; c_1}^{[a(t-1), b(t-1)]}, \dots, H, H_{c_1}, \dots), \quad (3.61)$$

$$r_{ab; c_1 \dots c_n \gamma}^{[a(s-1), b(s-1)]} = r_{ab; c_1 \dots c_n \gamma}^{[a(s-1), b(s-1)]}(R_{ab}^{[a(t-1), b(t-1)]}, R_{ab; c_1}^{[a(t-1), b(t-1)]}, \dots, H, H_{c_1}, \dots), \quad (3.62)$$

$$h_{c_1 \dots c_n \gamma} = h_{c_1 \dots c_n \gamma} (R_{ab}^{[a(t-1), b(t-1)]}, R_{ab; c_1}^{[a(t-1), b(t-1)]}, \dots, H, H_{c_1}, \dots), \quad (3.63)$$

$$dW^\alpha = 6 \frac{1}{2} \hat{f}_{\beta\gamma}^\alpha W^\beta \wedge W^\gamma, \quad (3.64)$$

$$dR_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]} = r_{ab; c_1 \dots c_n \gamma}^{[a(s-1), b(s-1)]} W^\gamma \Leftrightarrow \partial_\gamma R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]} = r_{ab; c_1 \dots c_n \gamma}^{[a(s-1), b(s-1)]}, \quad (3.65)$$

$$dH_{c_1 \dots c_n} = h_{c_1 \dots c_n \gamma} W^\gamma \Leftrightarrow \partial_\gamma H_{c_1 \dots c_n} = h_{c_1 \dots c_n \gamma}, \quad (3.66)$$

$$r_{ab; c_1 \dots c_n \gamma}^{[a(s-1), b(s-1)]} \frac{\partial \hat{f}_{\rho\sigma}^\alpha}{\partial R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}} + h_{c_1 \dots c_n \gamma} \frac{\partial \hat{f}_{\rho\sigma}^\alpha}{\partial H_{c_1 \dots c_n}} + \hat{f}_{\beta[\gamma}^\alpha \hat{f}_{\rho\sigma]}^\beta = 0, \quad (3.67)$$

$$r_{ab; c_1 \dots c_n [\rho}^{[a(s-1), b(s-1)]} \frac{\partial h_{\sigma]}]}{\partial R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}} + h_{c_1 \dots c_n [\rho} \frac{\partial h_{\sigma]}]}{\partial H_{c_1 \dots c_n}} + h_\alpha \hat{f}_{\rho\sigma}^\alpha = 0. \quad (3.68)$$

The input is $(\hat{f}_{\beta\gamma}^\alpha, h_\gamma)$, from which, all the rest equations are determined. The left hand sides of the $4d$ equations of motion (3.67)–(3.68) are polynomials of $(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$. For the randomly selected $(\hat{f}_{\beta\gamma}^\alpha, h_\gamma)$, (3.67)–(3.68) only has the trivial solution $R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]} = H_{c_1 \dots c_n} = 0$. A natural question is what might be the maximum on-shell degrees of freedom. If one can find such $(\hat{f}_{\beta\gamma}^\alpha, h_\gamma)$, for which, (3.67)–(3.68) is satisfied for the arbitrary $(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$, then there are no $4d$ equations of motion. This is not quite likely to be the case.⁷ By partially solving (3.67)–(3.68), one may determine $(\hat{f}_{\beta\gamma}^\alpha, h_\gamma)$, which, when plugged in (3.67)–(3.68), gives the $4d$ equations of motion for $(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$. In supergravity situation, the procedure is quite simple as is demonstrated in section 3.2. In higher spin theory, the more efficient way is to first determine the on-shell degrees of freedom $\Phi^{\tilde{\alpha}}$. Then with the off-shell $(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ expressed in terms of the on-shell $\Phi^{\tilde{\alpha}}$, we only need to find $(\hat{f}_{\beta\gamma}^\alpha, h_\gamma)$ satisfying the Bianchi identity for the arbitrary $\Phi^{\tilde{\alpha}}$. From the on-shell $(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ at one point, or the on-shell $(W_\mu^{[a(s-1), b(0)]}, H)$ in AdS_4 , (W^α, H) on \mathbf{M} can be determined via (3.64)–(3.66). With (W^α, H) on \mathbf{M} solved, the finite local higher spin transformation is the finite diffeomorphism transformation on \mathbf{M} , under which, $(W_\mu^{[a(s-1), b(0)]}, H)$ in one AdS_4 slice is moved to $(W_\mu^{[a(s-1), b(0)]}, H)$ in another AdS_4 slice. The higher spin symmetry is realized as an on-shell symmetry.

3.4 Relation with the unfolded equation in Vasiliev theory

With $\Phi^{\tilde{\alpha}}$ representing the on-shell degrees of freedom of $(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$, where $\tilde{\alpha}$ is in some representation of the Lorentz group, the unfolded equation becomes

$$\bar{f}_{\beta\gamma}^\alpha = \bar{f}_{\beta\gamma}^\alpha(\Phi^{\tilde{\sigma}}), \quad F_\gamma^{\tilde{\alpha}} = F_\gamma^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}}), \quad (3.69)$$

$$dW^\alpha = \frac{1}{2} \bar{f}_{\beta\gamma}^\alpha W^\beta \wedge W^\gamma, \quad d\Phi^{\tilde{\alpha}} = F_\beta^{\tilde{\alpha}} W^\beta, \quad (3.70)$$

$$F_{[\gamma}^{\tilde{\beta}} \frac{\partial \bar{f}_{\rho\sigma]}^\alpha}{\partial \Phi^{\tilde{\beta}}} + \bar{f}_{\beta[\gamma}^\alpha \bar{f}_{\rho\sigma]}^\beta = 0, \quad \frac{\partial F_{[\sigma}^{\tilde{\alpha}} F_{\rho]}^{\tilde{\beta}}}{\partial \Phi^{\tilde{\beta}}} + F_\gamma^{\tilde{\alpha}} \bar{f}_{\rho\sigma}^\gamma = 0. \quad (3.71)$$

⁷If it is true, then the $4d$ local HS gauge transformation (3.58) can be closed off-shell (for the arbitrary $W_\mu^{[a(s-1), b(0)]}$ and H in AdS_4).

It remains to find the suitable $(\bar{f}_{\beta\gamma}^\alpha, F_\gamma^\alpha)$ with the Bianchi identity (3.71) satisfied for the arbitrary $\Phi^{\tilde{\alpha}}$. Under the field redefinition $\Phi^{\tilde{\alpha}} \rightarrow \varphi^{\tilde{\alpha}} = \varphi^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}})$, $F_\beta^{\tilde{\alpha}} \rightarrow \frac{\partial \varphi^{\tilde{\alpha}}}{\partial \Phi^{\tilde{\sigma}}} F_\beta^{\tilde{\sigma}}$.

In Vasiliev theory, $\Phi^{\tilde{\alpha}} \sim \Phi^{[a(s+n), b(s)]}$ is in the twisted-adjoint representation of the higher spin algebra. (3.69) is obtained by solving the Vasiliev equation order by order. (3.71) is then automatically satisfied for the arbitrary $\Phi^{\tilde{\alpha}}$. Let us make a comparison between

$$\{H_{c_1 \dots c_n} | n = 0, 1, \dots\} \cup \{R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]} | s = 2, 4, \dots, n = 0, 1, \dots\} \quad (3.72)$$

and $\{\Phi^{[a(s+n), b(s)]}, s = 0, 2, \dots, n = 0, 1, \dots\}$. The two have the same number of indices, but the former is the off-shell field while the latter is on-shell. With the $4d$ equations of motion imposed on (3.72), the two may contain the same number of degrees of freedom.

Fields in the twisted-adjoint representation and the adjoint representation are related via the action of the Klein operator [6]

$$\Phi^\alpha t_\alpha \rightarrow \Phi^\alpha t_\alpha * \kappa = \Phi^\alpha \rho_\alpha^{\tilde{\alpha}} t_{\tilde{\alpha}} = \Phi^{\tilde{\alpha}} t_{\tilde{\alpha}}, \quad (3.73)$$

where $\Phi^{\tilde{\alpha}} = \rho_\alpha^{\tilde{\alpha}} \Phi^\alpha$, $\Phi^\alpha = (\rho^{-1})_\alpha^{\tilde{\alpha}} \Phi^{\tilde{\alpha}}$, $\rho_\alpha^{\tilde{\alpha}}$ is a constant matrix. For Φ in adjoint representation, i.e. $\Phi^\alpha \sim \Phi^{[a(s-1), b(t)]}$, it is possible to let $F_\beta^\alpha = \bar{f}_{\beta\gamma}^\alpha \Phi^\gamma$ [5], (3.69)–(3.71) reduces to

$$\bar{f}_{\beta\gamma}^\alpha = \bar{f}_{\beta\gamma}^\alpha(\Phi^\sigma), \quad (3.74)$$

$$dW^\alpha = \frac{1}{2} \bar{f}_{\beta\gamma}^\alpha W^\beta \wedge W^\gamma, \quad d\Phi^\alpha = \bar{f}_{\beta\gamma}^\alpha \Phi^\gamma W^\beta, \quad (3.75)$$

$$-\Phi^\nu \bar{f}_{\nu[\gamma}^\beta \frac{\partial \bar{f}_{\rho\sigma]}^\alpha}{\partial \Phi^\beta} + \bar{f}_{\beta[\gamma}^\alpha \bar{f}_{\rho\sigma]}^\beta = 0. \quad (3.76)$$

With $\Phi^\alpha \rightarrow \Phi^{\tilde{\alpha}}$, (3.69)–(3.71) is recovered for $F_\beta^{\tilde{\alpha}} = \bar{k}_{\beta\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}}$. $\bar{k}_{\beta\tilde{\gamma}}^{\tilde{\alpha}} = \rho_\alpha^{\tilde{\alpha}} \rho_\gamma^{\tilde{\gamma}} \bar{f}_{\beta\gamma}^\alpha$. $\Phi^{[a(0), b(0)]} \equiv \Phi = H$, $\partial_\beta \Phi = \partial_\beta H = H_\beta = \bar{k}_{\beta\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}}$. $\partial_\beta \Phi^{\tilde{\alpha}} = \bar{k}_{\beta\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}}$.

With (3.69)–(3.71) at hand, we have $R_{\beta\gamma}^\alpha = R_{\beta\gamma}^\alpha(\Phi^{\tilde{\sigma}}) = \bar{f}_{\beta\gamma}^\alpha(\Phi^{\tilde{\sigma}}) - f_{\beta\gamma}^\alpha$, $H_\beta = F_\beta(\Phi^{\tilde{\sigma}})$ since $d\Phi = F_\beta W^\beta$. Especially,

$$R_{ab}^{[a(s-1), b(s-1)]} = R_{ab}^{[a(s-1), b(s-1)]}(\Phi^{\tilde{\sigma}}), \quad H = \Phi, \quad (3.77)$$

and subsequently,

$$R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]} = R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}(\Phi^{\tilde{\sigma}}), \quad H_{c_1 \dots c_n} = H_{c_1 \dots c_n}(\Phi^{\tilde{\sigma}}), \quad (3.78)$$

where $\partial_c \Phi^{\tilde{\alpha}} = F_c^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}})$ is used.

$$\begin{aligned} \hat{f}_{\beta\gamma}^\alpha(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n}) &= \hat{f}_{\beta\gamma}^\alpha[R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}(\Phi^{\tilde{\sigma}}), H_{c_1 \dots c_n}(\Phi^{\tilde{\sigma}})] = \bar{f}_{\beta\gamma}^\alpha(\Phi^{\tilde{\sigma}}) \\ h_\gamma(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n}) &= h_\gamma[R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}(\Phi^{\tilde{\sigma}}), H_{c_1 \dots c_n}(\Phi^{\tilde{\sigma}})] = F_\gamma(\Phi^{\tilde{\sigma}}) \end{aligned} \quad (3.79)$$

Let us return to the discussion below (3.68). With $\hat{f}_{\beta\gamma}^\alpha$ and h_γ determined by the Bianchi identity, (3.67)–(3.68) may still have further constraints on $(R_{ab; c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$,

which are the $4d$ equations of motion. Alternatively, one may use the on-shell $\Phi^{\tilde{\sigma}}$ to parameterize the off-shell $(R_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ as is in (3.78). $4d$ equations of motion are then solved automatically. (3.67)–(3.68) does not impose any constraints on $\Phi^{\tilde{\sigma}}$. The key step in group manifold approach is to get the rheonomy condition and the $4d$ equations of motion from the Bianchi identity. For higher spin theory, the on-shell degrees of freedom form the twisted-adjoint representation of the higher spin algebra, while the Vasiliev equation gives an elegant way to solve the Bianchi identity. The solution for (W^α, H) in \mathbf{M} is characterized by the on-shell $(R_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ at one point, or by the arbitrary $\Phi^{\tilde{\alpha}}$ at that point. Nevertheless, it is $(R_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ that has the physical meaning. We are free to make a change of the variables $\varphi^{\tilde{\alpha}} = \varphi^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}})$ to use $\varphi^{\tilde{\alpha}}$ to parameterize $(R_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$. The good variables are those which are as relevant to $(R_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ as possible.

The nonlinear higher spin theory should also have the proper free theory limit that is equivalent to Fronsdal theory [18, 19]. In free theory limit, the equations of motion in (3.60)–(3.68) become

$$dW^\alpha - f_{(ab)\gamma}^\alpha W^{(ab)} \wedge W^\gamma - f_{a\gamma}^\alpha W^a \wedge W^\gamma = \frac{1}{2} R_{ab}^\alpha W^a \wedge W^b, \quad (3.80)$$

$$dR_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]} = r_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)](cd)} W^{(cd)} + R_{ab;c_1 \dots c_n c}^{[a(s-1), b(s-1)]} W^c, \quad (3.81)$$

$$dH_{c_1 \dots c_n} = h_{c_1 \dots c_n (cd)} W^{(cd)} + H_{c_1 \dots c_n c_{n+1}} W^{c_{n+1}}. \quad (3.82)$$

Since $r_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)](cd)} = \partial_{(cd)} R_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]}$ and $h_{c_1 \dots c_n (cd)} = \partial_{(cd)} H_{c_1 \dots c_n}$ give the local Lorentz transformation, (3.81)–(3.82) can be rewritten as

$$DR_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]} = R_{ab;c_1 \dots c_n c_{n+1}}^{[a(s-1), b(s-1)]} W^{c_{n+1}}, \quad (3.83)$$

$$DH_{c_1 \dots c_n} = H_{c_1 \dots c_n c_{n+1}} W^{c_{n+1}}, \quad (3.84)$$

where D is the standard covariant derivative with the connection $W^{(cd)}$. $DH = dH = H_c W^c$. For the theory to have the correct free theory limit, there will be

$$R_{ab}^{[a(s-1), b(t)]} = 0 \quad \text{for} \quad t \neq s-1 \quad (3.85)$$

so that (3.80) becomes

$$DW^{[a(s-1), b(t)]} = f_{a[c(s-1), d(t+1)]}^{[a(s-1), b(t)]} W^a \wedge W^{[c(s-1), d(t+1)]} + f_{a[c(s-1), d(t-1)]}^{[a(s-1), b(t)]} W^a \wedge W^{[c(s-1), d(t-1)]},$$

$$DW^{[a(s-1), b(s-1)]} = f_{a[c(s-1), d(s-2)]}^{[a(s-1), b(s-1)]} W^a \wedge W^{[c(s-1), d(s-2)]} + \frac{1}{2} R_{ab}^{[a(s-1), b(s-1)]} W^a \wedge W^b, \quad (3.86)$$

where $t < s-1$. (3.85) is also called the “central on-mass-shell theorem” [20, 21]. In Vasiliev theory, $R_{\beta\gamma}^\alpha$ satisfies (3.85) at the first order of the $\Phi^{\tilde{\alpha}}$ expansion.

Since the adjoint representation and the twisted-adjoint representation are related by a Klein transformation which is invertible, we may try to use Φ^α to parameterize $\bar{f}_{\beta\gamma}^\alpha$ as is in (3.74). If we further make a restriction that (3.75) can be written as

$$dW = H(W, \Phi), \quad d\Phi = F(W, \Phi) \quad (3.87)$$

with $H(W, \Phi)$ and $F(W, \Phi)$ polynomials of $W = W^\alpha t_\alpha$ and $\Phi = \Phi^\alpha t_\alpha$, the solution for (3.76) can be easily fixed, which is given in appendix C. Although the Bianchi identity is satisfied for the arbitrary Φ^α , (3.85) does not hold at the first order of the Φ^α expansion, so the theory does not have the correct free theory limit.

Satisfying the Bianchi identity for the on-shell $(R_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ and giving rise to the correct free theory limit are two requirements for $(\hat{f}_{\beta\gamma}^\alpha, h_\gamma)$. It is unclear whether the requirements can uniquely fix $(\hat{f}_{\beta\gamma}^\alpha, h_\gamma)$ or not. Starting from the rheonomy condition (3.14) in section 3.1, one may get (3.30) with no torsion constraint imposed on W_μ^α . The torsion constraint is just (3.85), or concretely,

$$R_{ab}^{[a(s-1), b(t)]} = g_{ab}^{[a(s-1), b(t)]} (W_\mu^\sigma, \partial_{\nu_1} W_\mu^\sigma, \dots, H, \partial_{\nu_1} H, \dots) = 0, \quad \text{for } t \neq s-1, \quad (3.88)$$

which will make W_μ^α reduce to $W_\mu^{[a(s-1), b(0)]}$ and also guarantee the correct free theory limit. In this case, having the correct free theory limit and satisfying the torsion constraint are the same thing. If there is such (3.14), for which the Bianchi identity on $(R_{ab;c_1 \dots c_n}^\alpha, H_{c_1 \dots c_n})$ reduces to the 4d equations of motion, then by setting $R_{ab}^{[a(s-1), b(t)]}$ to 0 for $t \neq s-1$, we will get (3.56) satisfying the Bianchi identity for the on-shell $(R_{ab;c_1 \dots c_n}^{[a(s-1), b(s-1)]}, H_{c_1 \dots c_n})$ and having the correct free theory limit. (3.85) holds exactly in this situation.

Finally, we will have a heuristic discussion on the group manifold approach to conformal HS theory. 3d conformal HS algebra and AdS₄ HS algebra are the same, so the corresponding group manifold is also **M**. The equations are still

$$dW^\alpha = \frac{1}{2} (f_{\beta\gamma}^\alpha + R_{\beta\gamma}^\alpha) W^\beta \wedge W^\gamma, \quad dH = H_\alpha W^\alpha. \quad (3.89)$$

The submanifold of interest is not AdS₄ but $\partial\text{AdS}_4 \subset \partial\mathbf{M}$. The solution of the unfolded equation in **M** is determined by the value of the 0-form multiplet at one point. In previous discussion, this point is selected at the bulk of AdS, but now, it should live at ∂AdS . The generated solution will remain at the near boundary region, since an infinite evolution is needed to move from the boundary to the bulk. The rheonomy condition $\bar{f}_{\beta\gamma}^\alpha = \bar{f}_{\beta\gamma}^\alpha(\Phi^{\tilde{\sigma}})$ and $F_\gamma^{\tilde{\alpha}} = F_\gamma^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}})$ in Vasiliev theory may undergo a reduction at the boundary with the role of $\Phi^{\tilde{\sigma}}$ played by a smaller set of 0-forms so that the solution at the near boundary region is determined by the dynamical fields in 3d.

In $\text{ho}(1|2 : [3, 2])$, the dilaton is $t_{0,4} = D$. It is convenient to choose the basis $\{t_\alpha\}$ with the definite conformal dimension, i.e. $[D, t_\alpha] = i\Delta_\alpha t_\alpha$. For example, the basis of $\text{so}(3, 2)$ is $\{D, P_i, K_i, L_{i,j}\}$ with $i, j = 1, 2, 3$. The dynamical fields are $W_m^{i_1 \dots i_{s-1}}$ in 3d, $m = 1, 2, 3$ [22]. Here $i_1 \dots i_{s-1}$ refers to the index of the spin s generator $P_{i_1 \dots i_{s-1}}$ with the dimension $1-s$.

There is a conjecture that the conformal HS theory at ∂AdS_{d+1} is related to the HS theory in AdS_{d+1} with the action of the conformal HS fields for even d equals to the logarithmically divergent term of the action of HS fields in AdS_{d+1} [23, 24]. In [15], the unfolded equation for a 3d conformal HS theory coming from the boundary limit of the AdS₄ Vasiliev theory was considered. It was shown that at ∂AdS_4 , $R_{ab}^\alpha \neq 0$ only when $t_\alpha = K_{i_1 \dots i_{s-1}}$. The condition could make W_m^α expressed in terms of the dynamical field $W_m^{i_1 \dots i_{s-1}}$ without imposing constraints on the latter. Correspondingly, in (3.56), the

independent 0-forms are $(R_{ij}^\alpha, R_{ij;k_1}^\alpha, \dots, H, H_{k_1}, \dots)$ for $t_\alpha = K_{i_1 \dots i_{s-1}}$, $s = 2, 4, \dots$. This is consistent with the fact that in odd dimensions, the conformal HS theory is trivial with no equations of motion imposed on dynamical fields [22, 24, 25].

On the other hand, in even dimensions, dynamical fields should satisfy Fradkin-Tseytlin equation [26]. The unfolded system of Fradkin-Tseytlin equation was formulated in [27, 28], where the 0-form multiplet is Weyl module generated by Weyl tensor, which, according to the terminology of [22], is the ground field strength. Equivalently, the 0-forms in (3.56) should now be taken as $(R_{ij}^{[i(s-1),j(s-1)]}, R_{ij;k_1}^{[i(s-1),j(s-1)]}, \dots, H, H_{k_1}, \dots)$. In free theory limit, $R_{ij}^\alpha = 0$ if $\Delta_\alpha < 0$, R_{ij}^α with $\Delta_\alpha > 0$ can all be expressed in terms of the derivatives of the Weyl tensor $R_{ij}^{[i(s-1),j(s-1)]}$. This is somewhat different from [15] for $3d$, where $R_{ij}^{[i(s-1),j(s-1)]} = 0$. It is interesting to consider the $4d$ conformal HS system arising from the boundary reduction of the Vasiliev equation in AdS_5 in analogy with [15]. In free theory limit, the obtained equation is expected to give the unfolded system of Fradkin-Tseytlin equation [27, 28]. The boundary value of the AdS_{d+1} HS fields was considered in [25, 29] in the ambient approach, where it was shown that for even d there is an obstruction for the bulk extension unless the conformal HS fields at ∂AdS_{d+1} satisfy the Fradkin-Tseytlin equation. In this case, the near boundary expansion of the on-shell AdS field (see, for example, [30]) does not have the logarithm term, which is required in the unfolded formalism, which in the minimal version does not allow for logarithmic terms to cancel the obstruction.

The unfolded equation for the $4d$ HS theory is invariant under the local Lorentz transformation $\text{SO}(3, 1)$, i.e. $R_{(a,b)\gamma}^\alpha = 0$. In [15], it is possible to impose the suitable boundary condition so that $R_{(0,4)i}^\alpha = 0$. If the conclusion can be extended to $R_{(0,4)\gamma}^\alpha = 0$, then the dilatation is unformed. Moreover, the original HS theory already have the undeformed $\text{SO}(3, 1)$ local Lorentz transformation, so by a naive counting, it seems that the inhomogeneous Weyl group \mathcal{IW} generated by $\{D, K_i, L_{i,j}\}$ can be undeformed at the boundary. In this case, the evolution along the $t_{0,4}$ direction is a conformal (gauge) transformation and the dynamics is reduced from $4d$ to $3d$. It remains to see whether there are consistent nonlinear unfolded equation for the conformal HS theory meeting this requirement. At least, the $3d$ local Lorentz transformation is undeformed.

3.5 The extended action principle for higher spin theory

In group manifold approach to supergravity, instead of imposing the rheonomy condition directly, one may construct the extended action whose variation gives both the rheonomy condition and the $4d$ equations of motion [3].

For example, in $\mathcal{N} = 1$ supergravity, the extended action is of the form

$$S = S[\nu^A, M_4] = \int_{M_4 \subset M} L^{(4)}(\nu^A), \quad (3.90)$$

where M_4 is a $4d$ submanifold of the superspace M ,⁸ and $L^{(4)}$ is a local Lorentz invariant 4-form in M constructed from ν^A via the exterior differentiation and the exterior product.

⁸We can use the group manifold \mathbf{M} instead of M , but the result is the same due to the factorization condition.

Variation of S with respect to both ν^A and M_4 gives

$$\frac{\delta L^{(4)}}{\delta \nu^A} = K_A^{(3)}(z) = 0. \quad (3.91)$$

$K_A^{(3)}$ is a 3-form that should vanish all over M . $K_A^{(3)}(z) = 0$ contains both the rheonomy condition and the $4d$ equations of motion. The concrete form of $L^{(4)}$ is

$$L^{(4)} = \epsilon_{abcd} R^{ab} \wedge \nu^c \wedge \nu^d + 4\bar{\psi} \wedge \gamma_5 \gamma_a \rho \wedge \nu^a, \quad (3.92)$$

where $\rho_{MN}^\alpha = R_{MN}^\alpha$.

For higher spin theory, if the extended action exists, it takes the form

$$S = S[W^\alpha, M_4] = \int_{M_4 \subset \mathbf{M}} L^{(4)}(W^\alpha), \quad (3.93)$$

where $L^{(4)}$ is a 4-form invariant under the local Lorentz transformation.

$$K_\sigma^{(3)} = \frac{\delta L^{(4)}}{\delta W^\sigma} = K_{\sigma[\alpha\beta\gamma]}^{(3)} W^\alpha \wedge W^\beta \wedge W^\gamma, \quad (3.94)$$

$$K_\sigma^{(3)} = 0 \Leftrightarrow K_{\sigma[\alpha\beta\gamma]}^{(3)} = 0. \quad (3.95)$$

We need to find the configuration W^α on \mathbf{M} with $K_\sigma^{(3)} = 0$ everywhere. Still, the on-shell solution on \mathbf{M} is characterized by the on-shell solution on M_4 . $M_4 \rightarrow M'_4$ is a diffeomorphism transformation on \mathbf{M} that is equivalent to the deformed higher spin gauge transformation. The equation $K_\sigma^{(3)} = 0$ is on-shell gauge invariant. Off-shell higher spin invariance has the further requirement $dL^{(4)} = 0$ [3]. Although the on-shell gauge invariance is automatically guaranteed, for the generic $L^{(4)}$, $K_\sigma^{(3)} = 0$ only has the trivial solution $R_{\beta\gamma}^\alpha = 0$, so the question is whether there is $L^{(4)}$ for which, the related $K_\sigma^{(3)} = 0$ has the nontrivial solution or not. In supergravity, having the nontrivial solution also puts the severe constraint on S .

In the simplest situation, if

$$L^{(4)} = \kappa_{\alpha\beta} R^\alpha \wedge R^\beta + \kappa_{\alpha\beta\gamma} R^\alpha \wedge W^\beta \wedge W^\gamma + \kappa_{\alpha\beta\gamma\sigma} W^\alpha \wedge W^\beta \wedge W^\gamma \wedge W^\sigma \quad (3.96)$$

with κ constants, then

$$K_{\sigma[\alpha\beta\gamma]}^{(3)} = -2\kappa_{\sigma\rho} f_{\chi[\alpha}^\rho R_{\beta\gamma]}^\chi - 2\kappa_{\rho\chi} f_{\sigma[\alpha}^\rho R_{\beta\gamma]}^\chi + \kappa_{\sigma\rho[\gamma} \hat{f}_{\alpha\beta]}^\rho + 2\kappa_{\rho\sigma[\gamma} R_{\alpha\beta]}^\rho + \kappa_{\rho[\beta\gamma} f_{\alpha]\sigma}^\rho + 4\kappa_{\sigma[\alpha\beta\gamma]}. \quad (3.97)$$

(3.97) imposes a set of linear relations among $R_{\beta\gamma}^\alpha$, which, when plugged into the Bianchi identity, may only allow the trivial solution $R_{\beta\gamma}^\alpha = 0$. The more general form of $L^{(4)}$ is

$$L^{(4)} = f_{\rho\sigma\chi\eta} (R_{\beta\gamma}^\alpha, \partial_\lambda R_{\beta\gamma}^\alpha, \dots) W^\rho \wedge W^\sigma \wedge W^\chi \wedge W^\eta \quad (3.98)$$

including an infinite number of derivatives. $K_{\sigma[\alpha\beta\gamma]}^{(3)} = 0$ are functions of $(R_{\beta\gamma}^\alpha, \partial_\lambda R_{\beta\gamma}^\alpha, \dots)$.

With $R_{\beta\gamma}^\alpha = \bar{f}_{\beta\gamma}^\alpha(\Phi^{\tilde{\sigma}}) - f_{\beta\gamma}^\alpha$ plugged in, $K_{\sigma[\alpha\beta\gamma]}^{(3)}$ should automatically vanish for the arbitrary $\Phi^{\tilde{\sigma}}$ if it is the action from which, the Vasiliev equation comes out. However, it is too complicated to fix the exact form of (3.98).

3.6 Dynamics of 0-form matter on group manifold with the fixed background

(3.60)–(3.68) describes the coupling of the spin 0 matter H and the spin $2, 4, \dots$ gravity field W^α . Under the local gauge transformation, which is the deformed higher spin transformation as well as the diffeomorphism transformation on \mathbf{M} , spin $0, 2, 4, \dots$ fields mix with each other. To describe the dynamics of the 0-form matter on \mathbf{M} with the fixed background, the matter-gravity coupling must be turned off. One may let $r_{\beta\gamma}^\alpha = 0$, then W_0^α gives the intrinsic geometry of the group manifold \mathbf{M} discussed in section 2. The equations of motion reduce to

$$dW_0^\alpha - \frac{1}{2}f_{\beta\gamma}^\alpha W_0^\beta \wedge W_0^\gamma = 0, \quad dH = H_\alpha W_0^\alpha \Leftrightarrow \partial_\alpha H = H_\alpha, \quad \partial_{[\rho} H_{\sigma]} + H_\alpha f_{\rho\sigma}^\alpha = 0. \quad (3.99)$$

The allowed gauge transformation parameter ϵ_0^α should satisfy

$$\delta_{\epsilon_0} W_{0\bar{M}}^\alpha = \partial_{\bar{M}} \epsilon_0^\alpha + f_{\beta\gamma}^\alpha \epsilon_0^\beta W_{0\bar{M}}^\gamma = 0, \quad (3.100)$$

generating the global higher spin transformation on \mathbf{M} .

$$\delta_{\epsilon_0} H = \xi_0^{\bar{M}} \partial_{\bar{M}} H = \epsilon_0^\beta H_\beta, \quad \delta_{\epsilon_0} H_\alpha = \xi_0^{\bar{M}} \partial_{\bar{M}} H_\alpha = \epsilon_0^\beta \partial_\beta H_\alpha. \quad (3.101)$$

(3.100) is integrable due to (3.99) with the solution characterized by ϵ_0^α at one point. With ϵ_0 satisfying (3.100), (3.99) is invariant under (3.101). $[\epsilon_0, \epsilon'_0]^\alpha = f_{\beta\gamma}^\alpha \epsilon_0^\beta \epsilon'^\gamma_0$. The structure constant is undeformed.

The next step is to impose the suitable rheonomy condition and derive the unfolded equation so that the solution on \mathbf{M} is determined by the (on-shell) fields in lower dimensions. In the following, we will consider two kinds of the rheonomy conditions which will make the final dynamics reduce to $4d$ and $3d$ respectively. The former gives a system equivalent to the linearized Vasiliev theory expanded on the background W_0^α , which also has an abelian local gauge symmetry invisible if we only focus on the equation for curvature. The latter comes from the $3d$ free massless scalar field theory at ∂AdS_4 . Since the $3d$ scalar forms the representation of the HS symmetry, it is possible to extend the scalar from $3d$ to (the boundary region of) \mathbf{M} with the global HS transformation realized as the isometry transformation.

3.6.1 The 4d global HS invariant system

Recall that in section 2, (2.22) and (2.25) are obtained. With $O(Z)$ replaced by $H(Z)$, from⁹

$$\begin{aligned} & \partial_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}} H \\ &= \sum_{r=0,2,\dots,s}^{t=1,2,\dots,2s+k-2r} G_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}}^{c_1\dots c_{2r+t}} \partial_{0,c_{2r+t}} \partial_{0,c_{2r+t-1}} \dots \partial_{0c_1\dots c_r, c_{r+1}\dots c_{2r+1}} H, \end{aligned} \quad (3.102)$$

$$\begin{aligned} & \partial_{0\dots 0a_1^1\dots a_{s_1}^1, b_1^1\dots b_{s_1+k_1}^1} \dots \partial_{0\dots 0a_1^p\dots a_{s_p}^p, b_1^p\dots b_{s_p+k_p}^p} H \\ & \sim \sum \Lambda(a_1\dots a_s, b_1\dots b_{s+k}) \partial_{0,b_{s+k}} \partial_{0,b_{s+k-1}} \dots \partial_{0a_1\dots a_s, b_1\dots b_{s+1}} H \end{aligned} \quad (3.103)$$

⁹(3.102) is the generic expansion, among which, some terms may vanish as is explained in appendix B.

for the constant G and Λ , the suitable rheonomy condition can be taken as

$$H_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}} = \sum_{r=0,2,\dots,s}^{t=1,2,\dots,2s+k-2r} G_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}}^{c_1\dots c_{2r+t}} H_{[0c_1\dots c_r, c_{r+1}\dots c_{2r+1}]; c_{2r+2}\dots c_{2r+t}} \quad (3.104)$$

for s even, k odd, r even; $H_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}} = 0$ for s odd, k even.

$$H_{[0c_1\dots c_r, c_{r+1}\dots c_{2r+1}]; c_{2r+2}\dots c_{2r+t}} = \partial_{c_{2r+t}} \dots \partial_{c_{2r+2}} H_{[0c_1\dots c_r, c_{r+1}\dots c_{2r+1}]} \quad (3.105)$$

According to the previous decomposition $\alpha = (A, Q)$, $\partial_Q H = H_Q = 0$, so

$$\partial_Q H_A = -f_{QA}^B H_B. \quad (3.106)$$

The evolution along the Q direction is a gauge transformation. The rest Bianchi identity is

$$\partial_A H_B = \partial_B H_A, \quad (3.107)$$

which is of course satisfied since (3.104) is obtained from the scalar operator $O(Z)$ on \mathbf{M} .

Based on (3.102) and (3.103), one may get the unfolded equation

$$\begin{aligned} \partial_\alpha H_{[0a_1\dots a_s, b_1\dots b_{s+1}]; c_1\dots c_n} &= h_{[0a_1\dots a_s, b_1\dots b_{s+1}]; c_1\dots c_n \alpha} \\ \partial_\alpha H_{c_1\dots c_n} &= h_{c_1\dots c_n \alpha} \\ \Leftrightarrow dH_{[0a_1\dots a_s, b_1\dots b_{s+1}]; c_1\dots c_n} &= h_{[0a_1\dots a_s, b_1\dots b_{s+1}]; c_1\dots c_n \alpha} W_0^\alpha \\ dH_{c_1\dots c_n} &= h_{c_1\dots c_n \alpha} W_0^\alpha, \end{aligned} \quad (3.108)$$

where $h_{[0a_1\dots a_s, b_1\dots b_{s+1}]; c_1\dots c_n \alpha}$ and $h_{c_1\dots c_n \alpha}$ are functions of $\{H_{c_1\dots c_n}, H_{[0a_1\dots a_s, b_1\dots b_{s+1}]; c_1\dots c_n} \mid n = 0, 1, \dots; s = 2, 4, \dots\}$. So the value of $(H_{[0a_1\dots a_s, b_1\dots b_{s+1}]; c_1\dots c_n}, H_{c_1\dots c_n})$ at one point determines its value on \mathbf{M} . Alternatively, $(H, H_{[0a_1 a_2, b_1 b_2 b_3]}, H_{[0a_1\dots a_4, b_1\dots b_5]}, \dots)$ on AdS_4 determines its value on \mathbf{M} .

The complete H_α is exhausted by $H_Q = 0$ and (3.104) for H_A . One may also add $H_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}}$ with s even, k even, which, together with H_A , forms the twisted-adjoint representation of the higher spin algebra. According to (B.10), for s even, $H_{[0a_1\dots a_s, b_1\dots b_{s+1}]}$ and $H_{[a_1\dots a_s, b_1\dots b_s]}$ are related via

$$H_{[0a_1\dots a_s, b_1\dots b_{s+1}]} = \sum_{\{b_1\dots b_{s+1}\}} \partial_{b_{s+1}} H_{[a_1\dots a_s, b_1\dots b_s]} + \dots \quad (3.109)$$

So $(H, H_{[0a_1 a_2, b_1 b_2 b_3]}, \dots)$ in AdS_4 is also equivalent to the field $(H, H_{[a_1 a_2, b_1 b_2]}, \dots)$, which is an irreducible representation of $G[\text{ho}(1|2 : [3, 2])]$.

The relation (3.104) is obtained from the operator $O(Z)$ on M . We may get the similar relation from the linearized Vasiliev theory, where $H_{[a_1\dots a_s, b_1\dots b_s]} \sim R_{a_1\dots a_s, b_1\dots b_s}^s$ gets the interpretation as the linearized curvature. Consider

$$dW_0^\alpha = \frac{1}{2} f_{\beta\gamma}^\alpha W_0^\beta \wedge W_0^\gamma \quad (3.110)$$

$$d\Phi^{\tilde{\alpha}} = k_{\beta\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}} W_0^\beta \Leftrightarrow \partial_\beta \Phi^{\tilde{\alpha}} = k_{\beta\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}} \quad (3.111)$$

$$d\tilde{W}^\alpha - f_{\beta\gamma}^\alpha W_0^\beta \wedge \tilde{W}^\gamma = \frac{1}{2} R_{1\beta\gamma}^\alpha (\Phi^{\tilde{\sigma}}) W_0^\beta \wedge W_0^\gamma \quad (3.112)$$

which is the linearized version of the Vasiliev equation (3.70) expanded on background W_0^α with \tilde{W}^α the fluctuation on it. $k_{\beta\tilde{\gamma}}^{\tilde{\alpha}} = \rho_{\alpha}^{\tilde{\alpha}} \rho_{\tilde{\gamma}}^{\gamma} f_{\beta\gamma}^{\alpha}$ is a constant. $k_{\beta\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}}$ is the lowest order term of $F_{\beta}^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}}) = k_{\beta\tilde{\gamma}}^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}}) \Phi^{\tilde{\gamma}}$ in (3.69). $R_{1\beta\gamma}^{\alpha}$ is the first order term of the polynomial $\tilde{f}_{\beta\gamma}^{\alpha}(\Phi^{\tilde{\sigma}})$ in (3.69), i.e. $\tilde{f}_{\beta\gamma}^{\alpha} = f_{\beta\gamma}^{\alpha} + R_{1\beta\gamma}^{\alpha} + \mathcal{O}(\Phi^2)$.

(3.110)–(3.112) are consistent if

$$\frac{\partial R_{1[\alpha\beta}^{\sigma}}{\partial \Phi^{\tilde{\rho}}} k_{\gamma]\tilde{\chi}}^{\tilde{\rho}} \Phi^{\tilde{\chi}} - f_{\rho[\alpha}^{\sigma} R_{1\beta\gamma]}^{\rho} - R_{1\rho[\alpha}^{\sigma} f_{\beta\gamma]}^{\rho} = 0 \quad (3.113)$$

$$k_{\sigma\tilde{\gamma}}^{\tilde{\alpha}} k_{\rho\tilde{\beta}}^{\tilde{\gamma}} - k_{\rho\tilde{\gamma}}^{\tilde{\alpha}} k_{\sigma\tilde{\beta}}^{\tilde{\gamma}} + f_{\rho\sigma}^{\chi} k_{\chi\tilde{\beta}}^{\tilde{\alpha}} = 0 \quad (3.114)$$

which is indeed the case due to the vanishing of the the first order part of the left hand side of (3.71).

(3.111)–(3.112) are invariant under the global HS transformation generated by $\xi_0^{\tilde{N}} = \epsilon_0^{\alpha} W_{0\alpha}^{\tilde{N}}$ preserving the background W_0^{α} .

$$d\epsilon_0^{\alpha} + f_{\beta\gamma}^{\alpha} \epsilon_0^{\beta} W_0^{\gamma} = 0 \quad (3.115)$$

as is in (3.100).

$$\begin{aligned} \delta W_0^{\alpha} &= 0, & \delta \Phi^{\tilde{\alpha}} &= \xi_0^{\tilde{N}} \partial_{\tilde{N}} \Phi^{\tilde{\alpha}} = k_{\beta\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}} \epsilon_0^{\beta}, \\ \delta \tilde{W}_M^{\alpha} &= \xi_0^{\tilde{N}} \partial_{\tilde{N}} \tilde{W}_M^{\alpha} + \partial_{\tilde{M}} \xi_0^{\tilde{N}} \tilde{W}_N^{\alpha}, & \delta R_{1\rho\sigma}^{\alpha} &= \xi_0^{\tilde{N}} \partial_{\tilde{N}} R_{1\rho\sigma}^{\alpha} = \frac{\partial R_{1\rho\sigma}^{\alpha}}{\partial \Phi^{\tilde{\alpha}}} k_{\beta\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}} \epsilon_0^{\beta}. \end{aligned} \quad (3.116)$$

There is also a residue local HS transformation

$$\delta \tilde{W}^{\alpha} = d\epsilon^{\alpha} + f_{\beta\gamma}^{\alpha} \epsilon^{\beta} W_0^{\gamma} = D\epsilon^{\alpha}, \quad \delta W_0^{\alpha} = 0, \quad \delta \Phi^{\tilde{\alpha}} = 0, \quad \delta R_{1\rho\sigma}^{\alpha} = 0, \quad (3.117)$$

which is invisible if we only focus on the equation for the 0-form. Intuitively, it seems that the global HS transformation for \tilde{W}^{α} and $R_{1\rho\sigma}^{\alpha}$ should be $\delta \tilde{W}^{\alpha} = f_{\beta\gamma}^{\alpha} \epsilon_0^{\beta} \tilde{W}^{\gamma}$ and $\delta R_{1\rho\sigma}^{\alpha} = f_{\beta\gamma}^{\alpha} \epsilon_0^{\beta} R_{1\rho\sigma}^{\gamma}$, which, however, is not consistent with the transformation law of $\Phi^{\tilde{\alpha}}$. The global HS transformation is a diffeomorphism transformation other than a gauge transformation.

Let us first consider (3.111). $H = \Phi$, $\partial_{\beta} H = \partial_{\beta} \Phi = k_{\beta\tilde{\gamma}} \Phi^{\tilde{\gamma}} = H_{\beta}$. In particular,

$$\partial_A H = \partial_A \Phi = k_{A\tilde{\gamma}} \Phi^{\tilde{\gamma}} = H_A = \Phi_A, \quad \partial_Q H = \partial_Q \Phi = k_{Q\tilde{\gamma}} \Phi^{\tilde{\gamma}} = H_Q = 0. \quad (3.118)$$

$$\begin{aligned} \partial_b \Phi^{[a_1 \dots a_{s+t}, b_1 \dots b_s]} &= k_b^{[a_1 \dots a_{s+t}, b_1 \dots b_s]}_{[c_1 \dots c_{s+t+1}, d_1 \dots d_s]} \Phi^{[c_1 \dots c_{s+t+1}, d_1 \dots d_s]} \\ &\quad + k_b^{[a_1 \dots a_{s+t}, b_1 \dots b_s]}_{[c_1 \dots c_{s+t-1}, d_1 \dots d_s]} \Phi^{[c_1 \dots c_{s+t-1}, d_1 \dots d_s]}. \end{aligned} \quad (3.119)$$

From (3.119), we have

$$\begin{aligned} \partial_b \Phi^{[a_1 \dots a_s, b_1 \dots b_s]} &= k_b^{[a_1 \dots a_s, b_1 \dots b_s]}_{[c_1 \dots c_{s+1}, d_1 \dots d_s]} \Phi^{[c_1 \dots c_{s+1}, d_1 \dots d_s]}, \\ \partial_b \Phi^{[a_1 \dots a_{s+1}, b_1 \dots b_s]} &= k_b^{[a_1 \dots a_{s+1}, b_1 \dots b_s]}_{[c_1 \dots c_{s+2}, d_1 \dots d_s]} \Phi^{[c_1 \dots c_{s+2}, d_1 \dots d_s]} + k_b^{[a_1 \dots a_{s+1}, b_1 \dots b_s]}_{[c_1 \dots c_s, d_1 \dots d_s]} \Phi^{[c_1 \dots c_s, d_1 \dots d_s]}, \\ \partial_b \Phi^{[a_1 \dots a_{s+2}, b_1 \dots b_s]} &= k_b^{[a_1 \dots a_{s+2}, b_1 \dots b_s]}_{[c_1 \dots c_{s+3}, d_1 \dots d_s]} \Phi^{[c_1 \dots c_{s+3}, d_1 \dots d_s]} + k_b^{[a_1 \dots a_{s+2}, b_1 \dots b_s]}_{[c_1 \dots c_{s+1}, d_1 \dots d_s]} \Phi^{[c_1 \dots c_{s+1}, d_1 \dots d_s]}, \\ &\dots \dots \dots \end{aligned} \quad (3.120)$$

so

$$\begin{aligned}
 \partial^{[a(s+1),b(s)]}\Phi &= \Phi^{[a(s+1),b(s)]} \sim \partial^b \Phi^{[a(s),b(s)]}, \\
 \partial^{[a(s+2),b(s)]}\Phi &= \Phi^{[a(s+2),b(s)]} \sim \partial^b \partial^b \Phi^{[a(s),b(s)]} + \Phi^{[a(s),b(s)]}, \\
 \partial^{[a(s+3),b(s)]}\Phi &= \Phi^{[a(s+3),b(s)]} \sim \partial^b \partial^b \partial^b \Phi^{[a(s),b(s)]} + \partial^b \Phi^{[a(s),b(s)]}, \\
 &\dots\dots
 \end{aligned} \tag{3.121}$$

Compared with the previous discussion on H_α , $H_{[a_1\dots a_s, b_1\dots b_s]}$ can be identified with $\Phi^{[a(s),b(s)]}$. (3.121) is also obtained in [12] by considering the 0-th level unfolded equation of Vasiliev theory, which is just (3.111).

In the interacting theory, $\partial_\beta \Phi^{\tilde{\alpha}} = \hat{k}_{\beta\tilde{\gamma}}^{\tilde{\alpha}}(\Phi^{\tilde{\sigma}})\Phi^{\tilde{\gamma}}$,

$$\begin{aligned}
 \partial_{b_1}\Phi^{[a(s),b(s)]} &= \hat{k}_{b_1\tilde{\gamma}}^{[a(s),b(s)]}\Phi^{\tilde{\gamma}}, \\
 \partial_{b_2}\partial_{b_1}\Phi^{[a(s),b(s)]} &= \frac{\partial \hat{k}_{b_1\tilde{\gamma}}^{[a(s),b(s)]}}{\partial \Phi^{\tilde{\sigma}}} \hat{k}_{b_2\tilde{\rho}}^{\tilde{\sigma}} \Phi^{\tilde{\rho}} \Phi^{\tilde{\gamma}} + \hat{k}_{b_1\tilde{\gamma}}^{[a(s),b(s)]} \hat{k}_{b_2\tilde{\rho}}^{\tilde{\gamma}} \Phi^{\tilde{\rho}}, \\
 &\dots\dots
 \end{aligned} \tag{3.122}$$

From (3.122), $\{\Phi^{\tilde{\alpha}} \sim \Phi^{[a(s),b(s+k)]}\}$ can be expressed in terms of $\{\partial_{b_k} \dots \partial_{b_1} \Phi^{[a(s),b(s)]}\}$, in a complicated way.

The interpretation of $\Phi^{[a(s),b(s)]}$ as the linearized curvature comes from (3.112). For the background geometry W_0^α in \mathbf{M} , one can always choose a particular gauge so that in AdS_4 ,

$$\{W_{0\mu}^\alpha\} = \{W_{0\mu}^a, W_{0\mu}^{(ab)}, 0, 0, \dots\}, \tag{3.123}$$

where $W_{0\mu}^a$ and $W_{0\mu}^{(ab)}$ are the vielbein and the connection characterizing AdS_4 geometry. (3.112) becomes (3.86) with $R_{ab}^{[a(s-1),b(s-1)]} \sim \Phi^{[a(s),b(s)]}$ the linearized Weyl tensor in free higher spin theory. (3.110)–(3.112) indicates that not only the interacting HS theory (Vasiliev theory), the free HS theory (Fronsdal theory) can also be consistently extended to \mathbf{M} with the symmetry reducing to the global HS symmetry and an abelian local HS symmetry.

Fronsdal equation for metric-like fields is invariant under the local HS transformation. A natural question is whether there are any manifestations of the global HS symmetry. Note that the “central on-mass-shell theorem” is the necessary condition for the linearized Vasiliev equation to reduce to the Fronsdal equation. In Vasiliev theory,

$$R_{1ab}^{[a(s-1),b(t)]}(\Phi^{\tilde{\sigma}}) = 0 \quad \text{for} \quad t \neq s-1 \tag{3.124}$$

is valid as a function equation independent of the position in \mathbf{M} . So the “central on-mass-shell theorem” is preserved under the global HS transformation, which is just a diffeomorphism transformation on \mathbf{M} . Under the global HS transformation, we move from one AdS fiber to another, with the Fronsdal equation satisfied as well. However, the transformation is on-shell and nontrivial, since we must first solve \tilde{W}^α all over \mathbf{M} and then perform (3.116). More explicitly, the transformation law of \tilde{W}_μ^α in AdS_4 is

$$\delta \tilde{W}_\mu^\alpha = \xi_0^{\tilde{N}} \partial_{\tilde{N}} \tilde{W}_\mu^\alpha + \partial_\mu \xi_0^{\tilde{N}} \tilde{W}_{\tilde{N}}^\alpha. \tag{3.125}$$

Suppose $\tilde{N} = \{\tilde{N}, \mu\}$, for simplicity, we may let $\tilde{W}_N^\alpha = 0$ in AdS_4 , then from (3.112),

$$\begin{aligned} \delta \tilde{W}_\mu^{[a(s-1), b(0)]} &= \xi_0^{\tilde{N}} W_{0\tilde{N}}^\beta \left(f_{\beta\gamma}^{[a(s-1), b(0)]} \tilde{W}_\mu^\gamma - R_{1a\beta}^{[a(s-1), b(0)]} W_{0\mu}^a \right) \\ &\quad + \xi_0^\nu \partial_\nu \tilde{W}_\mu^{[a(s-1), b(0)]} + \partial_\mu \xi_0^\nu \tilde{W}_\nu^{[a(s-1), b(0)]}. \end{aligned} \quad (3.126)$$

$R_{1a\beta}^\alpha(\Phi^{\tilde{\sigma}}) = r_{a\beta}^\alpha|_{\tilde{\sigma}} \Phi^{\tilde{\sigma}}$ with $r_{\beta\gamma}^\alpha|_{\tilde{\sigma}}$ the constant. $\Phi^{\tilde{\sigma}}$ can be expressed in terms of the $4d$ derivatives of $\Phi^{[a(s), b(s)]}$, which, in turn, is determined by \tilde{W}_μ^α and thus $\tilde{W}_\mu^{[a(s-1), b(0)]}$. The right hand side of (3.126) is a complicated $4d$ linear differential operator on $\tilde{W}_\mu^{[a(s-1), b(0)]}$. However, $\delta \tilde{W}_N^{[a(s-1), b(0)]} \neq 0$, the simplification condition $\tilde{W}_N^\alpha = 0$ is not preserved. In contrast to the gauge field, the global HS transformation law of the Weyl tensor in Fronsdal theory is straightforward.

$$\delta \Phi^{[a(s), b(s)]} = k_{\beta\tilde{\gamma}}^{[a(s), b(s)]} \Phi^{\tilde{\gamma}} \epsilon_0^\beta, \quad (3.127)$$

where $\Phi^{\tilde{\gamma}}$ can be written in terms of the $4d$ derivatives of $\Phi^{[a(s), b(s)]}$ via the relation $D_\mu \Phi^{\tilde{\alpha}} = k_{a\tilde{\gamma}}^{\tilde{\alpha}} \Phi^{\tilde{\gamma}} W_{0\mu}^a$.

It is well-known that the linearized Vasiliev theory is global HS invariant. By extending the space from AdS_4 to \mathbf{M} , the linearized Weyl module of the free higher spin theory can be compactly interpreted as $\partial_\alpha H$, the outer derivatives of a single scalar field H on \mathbf{M} .

3.6.2 The 3d global HS invariant system

The above $4d$ global HS invariant theory also has a local gauge symmetry. The genuine global HS invariant system without the local gauge symmetry is the $3d$ massless free scalar field theory living at ∂AdS_4 . In $3d$ free CFT, let ϕ be the operator for the dimension $1/2$ massless scalar and consider $\phi(X) = g(X)\phi(0')g(X)^{-1}$, $\forall g(X) \in G[\text{ho}(1|2 : [3, 2])]$. In contrast to $O(0)$ in the bulk, $\phi(0')$ is at the origin of ∂AdS_4 , so for the finite X , $\phi(X)$ is still at the near boundary region of \mathbf{M} with X the coordinate.

Scalar field at the near boundary region of \mathbf{M} also forms the representation of $G[\text{ho}(1|2 : [3, 2])]$. $-i\partial_\alpha \phi(X) = [Q_\alpha(X), \phi(X)]$. For $Q \in \text{so}(3, 2)$,

$$\begin{aligned} [K_m(0'), \phi(0')] &= 0, & [P_m(0'), \phi(0')] &= -i\partial_m \phi(0'), \\ [Q_{m,n}(0'), \phi(0')] &= 0, & [Q_{0,4}(0'), \phi(0')] &= -\frac{i}{2}\phi(0'). \end{aligned} \quad (3.128)$$

Generically, in $3d$ free CFT of the scalar ϕ , we have the relation

$$\partial_\alpha \phi(0') = i[Q_\alpha(0'), \phi(0')] = \sum_k (-i)^k \rho_\alpha^{i_1 \dots i_k} [P_{i_1}(0'), \dots [P_{i_k}(0'), \phi(0')] \dots], \quad (3.129)$$

where $i_k = 1, 2, 3$, ρ is the constant, because $\text{ho}(1|2 : [3, 2])$ can be realized as the quotient of the enveloping algebra of $\text{so}(3, 2)$ [13] (for HS algebra of any classical Lie algebras and in particular, \mathfrak{sp}_{2N} , \mathfrak{so}_N and \mathfrak{sl}_N , see [14]). As a result, the relation

$$\partial_\alpha \phi(X) = \sum_k (-i)^k \rho_\alpha^{i_1 \dots i_k} [P_{i_1}(X), \dots [P_{i_k}(X), \phi(X)] \dots] = \sum_k \rho_\alpha^{i_1 \dots i_k} \partial_{i_k} \dots \partial_{i_1} \phi(X) \quad (3.130)$$

is valid everywhere at the boundary region of \mathbf{M} for the constant ρ . $3d$ equations of motion for ϕ are also implicitly imposed by (3.130). The derivatives of ϕ in outer space can be expressed in terms of the derivatives of ϕ in inner space (∂AdS_4). This is not possible in (3.102), because the scalar field in AdS_4 cannot form the representation of the HS symmetry. One must introduce the higher spin fields, which, in (3.102), is reflected by $\partial_{0c_1 \dots c_r, c_{r+1} \dots c_{2r+1}}$.

Return to (3.99)–(3.101) and restrict to the near boundary region with H replaced by ϕ . From the rheonomy condition

$$\phi_\alpha = \sum_k \rho_\alpha^{i_1 \dots i_k} \partial_{i_k} \dots \partial_{i_1} \phi = \sum_k \rho_\alpha^{i_1 \dots i_k} \phi_{i_1 \dots i_k}, \quad (3.131)$$

one may get the unfolded equation

$$\partial_\alpha \phi_{i_1 \dots i_n} = \phi_{i_1 \dots i_n \alpha} \Leftrightarrow d\phi_{i_1 \dots i_n} = \phi_{i_1 \dots i_n \alpha} W_0^\alpha, \quad (3.132)$$

where $\phi_{i_1 \dots i_n \alpha}$ is the linear combination of $\{\phi, \phi_{i_1}, \phi_{i_1 i_2}, \dots\}$ with the constant coefficients.

The Bianchi identity

$$\frac{\partial \phi_{i_1 \dots i_k \gamma}}{\partial \phi_{j_1 \dots j_n}} \phi_{j_1 \dots j_n \beta} - \frac{\partial \phi_{i_1 \dots i_k \beta}}{\partial \phi_{j_1 \dots j_n}} \phi_{j_1 \dots j_n \gamma} + f_{\beta \gamma}^\alpha \phi_{i_1 \dots i_k \alpha} = 0 \quad (3.133)$$

is satisfied. From the on-shell $\{\phi, \phi_{i_1}, \phi_{i_1 i_2}, \dots\}$ at one point, or equivalent, the on-shell ϕ in ∂AdS_4 , ϕ in the near boundary region of \mathbf{M} can be determined. (3.132) is invariant under the global HS transformation

$$\delta_{\epsilon_0} \phi_{i_1 \dots i_n} = \xi_0^{\bar{N}} \partial_{\bar{N}} \phi_{i_1 \dots i_n} = \epsilon_0^\alpha \phi_{i_1 \dots i_n \alpha}. \quad (3.134)$$

In conclusion, to construct a theory with the global HS symmetry, we may try to find an unfolded equation like (3.108) and (3.132) for a 0-form multiplet on \mathbf{M} with the background geometry W_0^α . The equation should be integrable with the only dependence on \mathbf{M} comes from the 0-form and the 1-form W_0^α . Therefore, it is of course diffeomorphism invariant. The global higher spin transformation is a special diffeomorphism transformation preserving W_0^α .

4 Discussion

In supergravity, the rheonomy condition is simply $R_{BC}^A = r_{BC}^A(R_{ab}^{cd}, R_{ab}^\alpha, H, H_\alpha)$. Nevertheless, the most generic rheonomy condition in group manifold approach takes the form of (1.4) and (3.38) with all orders of derivatives included. If we make a similar truncation $R_{\beta\gamma}^\alpha = r_{\beta\gamma}^\alpha(R_{ab}^{[a(s-1), b(s-1)]}, H)$ in higher spin theory, then with $r_{\beta\gamma}^\alpha$ plugged into the Bianchi identity, we will get the $4d$ equations of motion, which, when expressed in terms of $(W_\mu^{[a(s-1), b(0)]}, H)$, do not contain derivatives higher than two. However, it is quite likely that such equations may only have the trivial solution $R_{ab}^{[a(s-1), b(s-1)]} = H = 0$ no matter how the coefficients in function $r_{\beta\gamma}^\alpha$ are adjusted. To allow for the nontrivial on-shell degrees of freedom, higher derivatives must be included so that $R_{\beta\gamma}^\alpha$ at one point is effectively determined by $(W_\mu^{[a(s-1), b(0)]}, H)$ on the whole AdS_4 . The $4d$ equations of motion for

$(W_\mu^{[a(s-1),b(0)]}, H)$ will also contain an infinite number of the higher derivative terms which make the theory nonlocal.

To write the unfolded equations (3.49) and (3.64)–(3.66), the infinite 0-form multiplets are necessarily involved in both supergravity and higher spin theory, since the solutions on the whole \mathbf{M} , including M_4/AdS_4 , are characterized by the on-shell 0-form multiplets at one point. For higher spin theory, the on-shell $(R_{ab;c_1\cdots c_n}^{[a(s-1),b(s-1)]}, H_{c_1\cdots c_n})$ is equivalent to $\{\Phi^{[a(s+n),b(s)]}\}$, so the solution on \mathbf{M} is also characterized by the arbitrary $\{\Phi^{[a(s+n),b(s)]}\}$ at that point. Merely based on group manifold approach without the knowledge of Vasiliev theory, we will finally arrive at (3.60)–(3.68) and then face the problem of finding the proper rheonomy condition that could solve the Bianchi identity, allow for the maximum on-shell degrees of freedom and have the correct free theory limit. It is Vasiliev theory that gives the solution meeting all these requirements. A question is whether there are other solutions or not. In appendix C, we give a rheonomy condition (for the bosonic higher spin theory) satisfying the Bianchi identity with the on-shell degrees of freedom $\{\Phi^\alpha\}$. However, the correct free theory limit is not recovered and the local Lorentz transformation is deformed.

In superspace with the fixed background geometry, the local super Poincaré symmetry reduces to the global super Poincaré symmetry. With the chiral constraint imposed, the component expansion of the scalar superfield in superspace gives the spin 0 and 1/2 fields (H, H^α) in M_4 . For higher spin theory, one can fix the background geometry of \mathbf{M} and then study the scalar field H in \mathbf{M} with the global higher spin symmetry. The component expansion of H gives the spin 0, 2, 4, \cdots fields $(H, H_{[a_1a_2,b_1b_2]}, H_{[a_1a_2a_3a_4,b_1b_2b_3b_4]}, \cdots)$ in AdS_4 , which, however, are not the gauge fields but the linearized Weyl tensors of the free HS theory, since the massless gauge fields are not the Lorentz tensor. Restricted to the near boundary region of \mathbf{M} , it is also possible to impose the rheonomy constraint so that the component expansion of H only gives the spin 0 field H in ∂AdS_4 . This is because although the $4d$ spin 0, 2, 4, \cdots fields all together form the representation of the HS symmetry, the $3d$ spin 0 field alone forms the HS representation.

Acknowledgments

This research was supported in part by the Natural Science Foundation of China under grant numbers 10821504, 11075194, 11135003, 11275246, and 11475238, and by the National Basic Research Program of China (973 Program) under grant number 2010CB833000.

A $G[\text{ho}(1|2 : [3, 2])]/E$ from the CFT operators

The minimal bosonic higher spin theory in AdS_4 is dual to the $3d$ $O(N)$ vector model [31, 32]. The conserved charges in both theories form the algebra isomorphic to $\text{ho}(1|2 : [3, 2])$. The duality requires that the states and the operators in CFT and the bulk theory can be identified, so in the following, we will use the operators in CFT to represent their counterparts in $4d$ HS theory. Suppose $\{Q_\alpha \sim Q_{A_1\cdots A_{s-1}, B_1\cdots B_{s-1}}\}$ are charge operators in CFT corresponding to $\{t_\alpha \sim t_{A_1\cdots A_{s-1}, B_1\cdots B_{s-1}}\}$ in (2.2). $A_k, B_k = 0, 1, 2, 3, 4$. The explicit form

of Q_α can be found in [33]. The CFT realization of the bulk local field operators is usually constructed as [34–36]

$$\Phi(x) \leftrightarrow \int dX K(X|x) \mathcal{O}(X) \quad (\text{A.1})$$

in large N limit, where $\Phi(x)$ is the bulk field in AdS, $\mathcal{O}(X)$ is the boundary operator in CFT, $K(X|x)$ is the boundary-bulk propagator. $\Phi(x)$ like this of course satisfies the free field equation in AdS, which is acceptable when $N \rightarrow \infty$. In this section and the next one, we will construct the spin 0 field operator and the spin s linearized curvature tensor operators in AdS for $s = 2, 4, \dots$, using the CFT operators $O_{i_1 \dots i_s}$ in [33]. $i_k = 1, 2, 3$. $O_{i_1 \dots i_s}(x)$ only contains the positive frequency part, but it is enough for the present use.

For AdS_4 parameterized by $x_0^2 - x_1^2 - x_2^2 - x_3^2 + x_4^2 = 1$, let $Q_{A,B}$ be the generators of $\text{SO}(3, 2)$, then for an operator $O(x)$ with $s = 0$,

$$[Q_{A,B}, O(x)] = i(x_A \partial_B - x_B \partial_A) O(x). \quad (\text{A.2})$$

Without losing of the generality, let us consider a point 0 in the bulk of AdS_4 with the coordinate $x^0 = 1, x^1 = \dots = x^4 = 0$.

$$[Q_{m,n}, O(0)] = [Q_{m,4}, O(0)] = [K_m - iP_m, O(0)] = 0, \quad (\text{A.3})$$

where $m, n = 1, 2, 3$. $\{Q_{0,4}, Q_{0,m}\} \subset K(0)$ generates the tangent space along AdS_4 . From (A.3), according to the operators constructed in [33], $O(0)$ is solved as

$$O(0) = \sum \frac{k!}{(2k+1)!!} a_{i_1 \dots i_k}^+ a_{i_1 \dots i_k}^+. \quad (\text{A.4})$$

In [33], the generic elements of $\text{ho}(1|2 : [3, 2])$ can be written as

$$Q_{m_1 \dots m_p, n_1 \dots n_q} = i \sum g(l) a_{m_1 \dots m_p i_1 \dots i_l}^+ a_{n_1 \dots n_q i_1 \dots i_l} \quad (\text{A.5})$$

with $m_k, n_k, i_k = 1, 2, 3$, so

$$[Q_{m_1 \dots m_p, n_1 \dots n_q}, O(0)] \sim i \sum g(l) a_{m_1 \dots m_p i_1 \dots i_l}^+ a_{n_1 \dots n_q i_1 \dots i_l}^+. \quad (\text{A.6})$$

One can choose the basis $\{Q\}$ of $\text{ho}(1|2 : [3, 2])$ with the definite conformal dimension.

$$[D, Q] = -i\Delta Q, \quad [D, Q^+] = i\Delta Q^+. \quad (\text{A.7})$$

Let $H_Q = Q + Q^+$, $\bar{H}_Q = i(Q - Q^+)$, there will be

$$[H_Q, O(0)] = 0, \quad [\bar{H}_Q, O(0)] = 2i[Q, O(0)]. \quad (\text{A.8})$$

$a[E(0)] = \{H_Q\}$, $K(0) = \{\bar{H}_Q\}$. Moreover,

$$[\{\bar{H}_Q\}, \{\bar{H}_Q\}] \subset \{H_Q\}, \quad [\{H_Q\}, \{H_Q\}] \subset \{H_Q\}, \quad [\{H_Q\}, \{\bar{H}_Q\}] \subset \{\bar{H}_Q\}. \quad (\text{A.9})$$

M is a symmetric space.

For $\text{ho}(1|2 : [3, 2])$, there is an involution σ

$$\sigma(Q) = Q^+ \quad (\text{A.10})$$

with $\sigma^2 = 1$. σ has the eigenvalues 1 and -1 with $\{H_Q\}$ and $\{\bar{H}_Q\}$ defined above the corresponding eigenspaces.

$$\text{ho}(1|2 : [3, 2]) = \{H_Q\} \oplus \{\bar{H}_Q\}. \quad (\text{A.11})$$

Under the Wick rotation, $x^0 \rightarrow ix^0$, the action of σ is then $\sigma : ix^0 \rightarrow -ix^0$, so

$$\begin{aligned} a[E(0)] &= \{H_Q\} = \{t_{0\dots 0a_1, b_1\dots b_{s-1}}, t_{0\dots 0a_1a_2a_3, b_1\dots b_{s-1}}, \dots, t_{a_1\dots a_{s-1}, b_1\dots b_{s-1}}\}, \\ K(0) &= \{\bar{H}_Q\} = \{t_{0\dots 0, b_1\dots b_{s-1}}, t_{0\dots 0a_1a_2, b_1\dots b_{s-1}}, \dots, t_{0a_1\dots a_{s-2}, b_1\dots b_{s-1}}\}. \end{aligned} \quad (\text{A.12})$$

The decomposition is consistent with (2.3) and (2.4).

B CFT realization of the spin s linearized Riemann tensor operator in AdS_4

In radial quantization of the $3d$ $O(N)$ vector model, for each $s = 0, 2, \dots$, there is a unique primary operator $O_{i_1\dots i_s}(0')$ with spin s .

$$\frac{1}{2}[Q^{A,B}, [Q_{A,B}, O_{i_1\dots i_s}(0')]] \equiv [C_2, O_{i_1\dots i_s}(0')] = 2(s^2 - 1)O_{i_1\dots i_s}(0'), \quad (\text{B.1})$$

where C_2 is the Casimir operator. $i_k = 1, 2, 3$. Here $0'$ represents the origin in the boundary CFT and should be distinguished from the 0 in appendix A. $\{\partial_{\mu_1} \dots \partial_{\mu_n} O_{i_1\dots i_s}(0') | s = 0, 2, \dots; n = 0, 1, \dots\}$ forms the complete basis of the 1-particle Hilbert space. $\mu_k = 1, 2, 3$. $\{O(0'), O_{i_1 i_2}(0'), \dots\}$ is the higher spin multiplet. The action of the generic Q_α on the spin 0 primary operator $O(0') = a^+ a^+$ can be decomposed as

$$[Q_{A_1\dots A_{s-1}, B_1\dots B_{s-1}}, O(0')] = \sum_{r=0,2,\dots}^{t=0,1,\dots} g_{A_1\dots A_{s-1}, B_1\dots B_{s-1}}^{\mu_1\dots\mu_t; i_1\dots i_r} \partial_{\mu_1} \dots \partial_{\mu_t} O_{i_1\dots i_r}(0'). \quad (\text{B.2})$$

Let us construct the $\text{SO}(3, 1)$ tensor operator with spin s in the sense of (B.1) in AdS bulk. Such operator does not represent the spin s gauge field which is not a tensor, but rather the field strength of it. The spin 0 operator $O(0)$ is already given by (A.4). For operators with the higher spin, consider

$$\begin{aligned} A_{m_1 m_2, k}^+ &= a_{m_1 m_2 i_1 \dots i_k}^+ a_{i_1 \dots i_k}^+ + f_1 a_{m_1 i_1 \dots i_k}^+ a_{m_2 i_1 \dots i_k}^+, \\ A_{m_1 m_2 m_3 m_4, k}^+ &= a_{m_1 m_2 m_3 m_4 i_1 \dots i_k}^+ a_{i_1 \dots i_k}^+ + f_1 a_{m_1 m_2 m_3 i_1 \dots i_k}^+ a_{m_4 i_1 \dots i_k}^+ + \dots \\ &\quad + f_4 a_{m_2 m_3 m_4 i_1 \dots i_k}^+ a_{m_1 i_1 \dots i_k}^+ + f_5 a_{m_1 m_2 i_1 \dots i_k}^+ a_{m_3 m_4 i_1 \dots i_k}^+ \\ &\quad + f_6 a_{m_1 m_3 i_1 \dots i_k}^+ a_{m_2 m_4 i_1 \dots i_k}^+ + f_7 a_{m_1 m_4 i_1 \dots i_k}^+ a_{m_2 m_3 i_1 \dots i_k}^+, \\ &\quad \dots \end{aligned} \quad (\text{B.3})$$

which is the most generic s -tensor with the dimension $s + 2k + 1$. $m_p, i_p = 1, 2, 3$. For each s , imposing the condition

$$[C_2, A_{m_1 \dots m_s, k}^+] = 2(s^2 - 1)A_{m_1 \dots m_s, k}^+ \quad (\text{B.4})$$

can uniquely fix f_i in (B.3). The corresponding operator is denoted as $A_{m_1 \dots m_s, k}^{(s)+}$, which is totally symmetric and traceless.

Suppose

$$O_{m_1 \dots m_s}^s(0) = \sum g(k) A_{m_1 \dots m_s, k}^{(s)+} \quad (\text{B.5})$$

is a spin s tensor operator at 0, then in analogy with (A.3),¹⁰

$$[Q_{m, n}, O_{m_1 \dots m_s}^s(0)] = \Sigma_{mn} O_{m_1 \dots m_s}^s(0), \quad [Q_{4, m}, O_{m_1 \dots m_s}^s(0)] = \Sigma_{4m} O_{m_1 \dots m_s}^s(0), \quad (\text{B.6})$$

$m, n = 1, 2, 3$. Σ is the spin operator. The first equation in (B.6) is satisfied for the arbitrary $g(k)$. $O_{m_1 \dots m_s}^s(0)$ forms the representation of $\text{SO}(3)$. The complete $\text{SO}(3, 1)$ representation can be obtained by the successive action of $Q_{4, m}$. The coefficient $g(k)$ in (B.5) is determined by the requirement that at some point, no new operators can be created as is required by the second equation of (B.6). When $s = 0$, the solution of $[Q_{4, m}, O^0(0)] = 0$ is $O(0)$ in (A.4). When $s = 2$, the minimal times for the action of $Q_{4, m}$ is 3. The corresponding $O_{m_1 m_2}^2(0)$ can be written as $O_{m_1 m_2, 44}^2(0)$, while the action of $\{Q_{4, m}, Q_{m, n}\}$ gives the complete $\text{SO}(3, 1)$ representation $O_{b_1 b_2, b_3 b_4}^2(0)$ with $b_i = 1, 2, 3, 4$. Generically, for spin s operator $O_{m_1 \dots m_s}^s(0)$, we have $O_{m_1 \dots m_s}^s(0) \equiv O_{m_1 \dots m_s, 4 \dots 4}^s(0)$ with the $\text{SO}(3, 1)$ completion $O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}}^s(0)$. The maximum number of 4 in $O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}}^s(0)$ is s .

The minimum spin s $\text{SO}(3, 1)$ tensor operator transforming like (B.6) is not $O_{b_1 \dots b_s}^s$ but $O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}}^s$. This is expected, since the massless gauge field is not a Lorentz tensor. $O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}}^s$ matches well with the Riemann curvature $R_{b_1 \dots b_s, b_{s+1} \dots b_{2s}}^s$ of the spin s field, or more precisely, the linearized Riemann curvature in AdS background since $O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}}^s$ only creates single particle states.

$\{Q_{0, b}\}$ generates the tangent space at 0 along AdS_4 . The Successive action of $Q_{0, b}$ gives

$$O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}; b_{2s+1} \dots b_{2s+k}}^s(0) = [Q_{0, b_{2s+k}}, \dots [Q_{0, b_{2s+2}}, [Q_{0, b_{2s+1}}, O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}}^s(0)]] \dots]. \quad (\text{B.7})$$

$O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}; b_{2s+1} \dots b_{2s+k}}^s(0)$ is the descendant of $O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}}^s(0)$ thus is a spin s operator as well. $\forall x \in \text{AdS}_4$,¹¹

$$\begin{aligned} O_{b_1 \dots b_{2s}}^s(x) &= g(x) O_{b_1 \dots b_{2s}}^s(0) g(x)^{-1}, \\ O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}; b_{2s+1} \dots b_{2s+k}}^s(x) &= g(x) O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}; b_{2s+1} \dots b_{2s+k}}^s(0) g(x)^{-1}, \end{aligned} \quad (\text{B.8})$$

$g(x) \in \text{SO}(3, 2)$. $\{O_{b_1 \dots b_s, b_{s+1} \dots b_{2s}; b_{2s+1} \dots b_{2s+k}}^s(x) | s = 0, 2, \dots; k = 0, 1, \dots\}$ at x compose the complete basis for the 1-particle Hilbert space of the $4d$ HS theory.

¹⁰ $O_{m_1 \dots m_s}^s(0)$ is a gauge invariant operator. The $\text{SO}(3, 1)$ transformation of the spin s massless gauge field also has the derivative terms on the right hand side.

¹¹The relation (B.8) is not valid for $O_{i_1 \dots i_s}$, which is not a tensor.

Now consider $[Q_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}}, O(0)]$ with $k = 1, 3, \dots$, $a_i, b_i = 1, 2, 3, 4$, which could be expanded as

$$\begin{aligned}
 & [Q_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}}, O(0)] \\
 = & \sum_{r=0,2,\dots,s}^{t=0,1,\dots,2s+k-2r} k_{a_1\dots a_s, b_1\dots b_{s+k}}^{c_1\dots c_{2r+t}} [Q_{0,c_{2r+t}}, \dots [Q_{0,c_{2r+2}}, [Q_{0,c_{2r+1}}, O_{c_1\dots c_r, c_{r+1}\dots c_{2r}}^r(0)]] \dots].
 \end{aligned} \tag{B.9}$$

(B.2) could be taken as the boundary limit of (B.9), where the linearized curvature tensor has already been written as the derivatives of the metric operator $O_{i_1\dots i_s}$. There is no charge operator that could directly create $O_{a_1\dots a_s, b_1\dots b_s}^s(0)$ from $O(0)$, the closest one is

$$[Q_{0a_1\dots a_s, b_1\dots b_{s+1}}, O(0)] = \sum_{\{b_1\dots b_{s+1}\}} [Q_{0,b_{s+1}}, O_{a_1\dots a_s, b_1\dots b_s}^s(0)] + \dots \tag{B.10}$$

\dots are possible terms with the spin lower than s . (B.9) and (B.10) represent the generic possibilities, among which some terms may simply vanish. Since k is odd, instead of the “primary” $O_{a_1\dots a_s, b_1\dots b_s}^s(0)$, one can also use the less “primary” $[Q_{0a_1\dots a_s, b_1\dots b_{s+1}}, O(0)]$,

$$\begin{aligned}
 & [Q_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}}, O(0)] \\
 = & \sum_{r=0,2,\dots,s}^{t=1,\dots,2s+k-2r} f_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}}^{c_1\dots c_{2r+t}} [Q_{0,c_{2r+t}}, \dots [Q_{0,c_{2r+2}}, [Q_{0c_1\dots c_r, c_{r+1}\dots c_{2r+1}}, O(0)]] \dots] \\
 = & \sum_{r=0,2,\dots,s}^{t=1,\dots,2s+k-2r} g_{0\dots 0a_1\dots a_s, b_1\dots b_{s+k}}^{c_1\dots c_{2r+t}} [Q_{0c_1\dots c_r, c_{r+1}\dots c_{2r+1}}, \dots [Q_{0,c_{2r+t-1}}, [Q_{0,c_{2r+t}}, O(0)]] \dots].
 \end{aligned} \tag{B.11}$$

Especially,

$$[Q_{0\dots 0, b_1\dots b_k}, O(0)] \sim \sum_{\{b_1\dots b_k\}} [Q_{0,b_k}, \dots [Q_{0,b_2}, [Q_{0,b_1}, O(0)]] \dots], \tag{B.12}$$

$[Q_{0\dots 0, b_1\dots b_k}, O(0)]$ is the descendant of $O(0)$.

C A rheonomy condition satisfying the Bianchi identity without giving the correct free theory limit

For the 0-form in adjoint representation of $\text{ho}(1|2 : [3, 2])$, the equations of motion and the gauge transformation are

$$dW^\alpha = \frac{1}{2} \bar{f}_{\beta\gamma}^\alpha W^\beta \wedge W^\gamma, \quad d\Phi^\alpha = \bar{f}_{\beta\gamma}^\alpha W^\beta \Phi^\gamma \tag{C.1}$$

and

$$\delta_\epsilon W^\alpha = d\epsilon^\alpha + \bar{f}_{\beta\gamma}^\alpha \epsilon^\beta W^\gamma, \quad \delta_\epsilon \Phi^\alpha = \bar{f}_{\beta\gamma}^\alpha \epsilon^\beta \Phi^\gamma \tag{C.2}$$

with

$$\bar{f}_{\beta[\gamma}^{\alpha} \bar{f}_{\rho\sigma]}^{\beta} - \Phi^{\nu} \bar{f}_{\nu[\gamma}^{\beta} \frac{\partial \bar{f}_{\rho\sigma]}^{\alpha}}{\partial \Phi^{\beta}} = 0. \quad (\text{C.3})$$

Expanding in terms of Φ^{α} ,

$$\bar{f}_{\beta\gamma}^{\alpha} = f_{\beta\gamma}^{\alpha} + f_{\beta\gamma|\sigma_1}^{\alpha} \Phi^{\sigma_1} + f_{\beta\gamma|\sigma_1\sigma_2}^{\alpha} \Phi^{\sigma_1} \Phi^{\sigma_2} + \dots \quad (\text{C.4})$$

If we assume

$$t_{\alpha} f_{\beta\gamma|\sigma_1 \dots \sigma_n}^{\alpha} = f(t_{\beta}, t_{\gamma}; t_{\sigma_1} \dots t_{\sigma_n}), \quad (\text{C.5})$$

where $f(t_{\beta}, t_{\gamma}; t_{\sigma_1} \dots t_{\sigma_n})$ is the sum of the product of $t_{\beta}, t_{\gamma}, t_{\sigma_1}, \dots, t_{\sigma_n}$ with t_{β} and t_{γ} antisymmetric, $t_{\sigma_1} \dots t_{\sigma_n}$ symmetric, then

$$t_{\alpha} \bar{f}_{\beta\gamma}^{\alpha} = f(t_{\beta}, t_{\gamma}) + f(t_{\beta}, t_{\gamma}; \Phi) + f(t_{\beta}, t_{\gamma}; \Phi, \Phi) + \dots \quad (\text{C.6})$$

$\Phi = \Phi^{\alpha} t_{\alpha}$. With (C.4) plugged in (C.3), comparing the coefficients order by order, the solution can only be

$$t_{\alpha} \bar{f}_{\beta\gamma}^{\alpha} = [t_{\beta}, t_{\gamma}] F(\Phi), \quad \text{or} \quad t_{\alpha} \bar{f}_{\beta\gamma}^{\alpha} = F(\Phi) [t_{\beta}, t_{\gamma}], \quad (\text{C.7})$$

where $F(\Phi)$ is an arbitrary function of Φ with $F(0) = 1$. Plug (C.7) into (C.3), we can see (C.3) is indeed satisfied. (C.1) and (C.2) become

$$dW = [W, W] F(\Phi), \quad d\Phi = [W, \Phi] F(\Phi) \quad (\text{C.8})$$

and

$$\delta_{\epsilon} W = d\epsilon + [\epsilon, W] F(\Phi), \quad \delta_{\epsilon} \Phi = [\epsilon, \Phi] F(\Phi), \quad (\text{C.9})$$

or

$$dW = F(\Phi) [W, W], \quad d\Phi = F(\Phi) [W, \Phi] \quad (\text{C.10})$$

and

$$\delta_{\epsilon} W = d\epsilon + F(\Phi) [\epsilon, W], \quad \delta_{\epsilon} \Phi = F(\Phi) [\epsilon, \Phi]. \quad (\text{C.11})$$

$$\bar{f}_{\beta\gamma}^{\alpha} = \langle [t_{\beta}, t_{\gamma}] F(\Phi) | t^{\alpha} \rangle = f_{\beta\gamma}^{\sigma} \langle t_{\sigma} F(\Phi) | t^{\alpha} \rangle \quad \text{or} \quad \bar{f}_{\beta\gamma}^{\alpha} = \langle F(\Phi) [t_{\beta}, t_{\gamma}] | t^{\alpha} \rangle = f_{\beta\gamma}^{\sigma} \langle F(\Phi) t_{\sigma} | t^{\alpha} \rangle. \quad (\text{C.12})$$

Each $F(\Phi)$ gives a consistent deformation of $\bar{f}_{\beta\gamma}^{\alpha} = f_{\beta\gamma}^{\alpha}$. With the field redefinition $\Phi' = f(\Phi)$, (C.8)–(C.11) become

$$dW = [W, W] F[f^{-1}(\Phi')], \quad d\Phi' = [W, \Phi'] F[f^{-1}(\Phi')] \quad (\text{C.13})$$

and

$$\delta_{\epsilon} W = d\epsilon + [\epsilon, W] F[f^{-1}(\Phi')], \quad \delta_{\epsilon} \Phi' = [\epsilon, \Phi'] F[f^{-1}(\Phi')], \quad (\text{C.14})$$

or

$$dW = F[f^{-1}(\Phi')] [W, W], \quad d\Phi' = F[f^{-1}(\Phi')] [W, \Phi'] \quad (\text{C.15})$$

and

$$\delta_{\epsilon} W = d\epsilon + F[f^{-1}(\Phi')] [\epsilon, W], \quad \delta_{\epsilon} \Phi' = F[f^{-1}(\Phi')] [\epsilon, \Phi']. \quad (\text{C.16})$$

Especially, when $f = F$, $F[f^{-1}(\Phi)] = \Phi'$. All of the consistent deformations are related to $F(\Phi) = \Phi$ by a field redefinition.

Until now, we have not made any assumption on the algebra $\{t_\alpha\}$, so (C.12) holds for the arbitrary algebra which is also a ring. Consider the 4d bosonic higher spin theory with the spin $s = 0, 1, 2, \dots$ and the algebra g , for $t_\alpha \in g$, $t_\alpha \sim t_{a_1 \dots a_{s-1}, b_1 \dots b_t 0 \dots 0} \sim y^m \bar{y}^n$ with $m + n = 2(s - 1)$, $t = |m - n|/2$. $\forall t_\alpha, t_\beta \in g$, $t_\alpha t_\beta \in g$, so (C.8)–(C.11) are well defined, but the truncation to the minimal bosonic higher spin theory is not possible. The theory does not have the correct free theory limit since (3.85) is not satisfied. Also, $R_{(ab)\gamma}^\alpha \neq 0$, the local Lorentz transformation is deformed.

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